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Institut de Ciències del Cosmos  
UNIVERSITAT DE BARCELONA



# The neutron skin thickness in atomic nuclei

Xavier Roca-Maza  
University of Tokyo

April 15<sup>th</sup> 2026

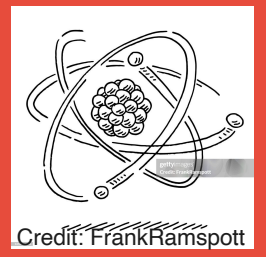


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# Neutron skin thickness in atomic nuclei

What is it?



The **mean square radius** of the **neutron** (n) or **proton** (p) **density distribution** ( $\rho_q(\vec{r})$  with  $q = n, p$ ) is defined as:

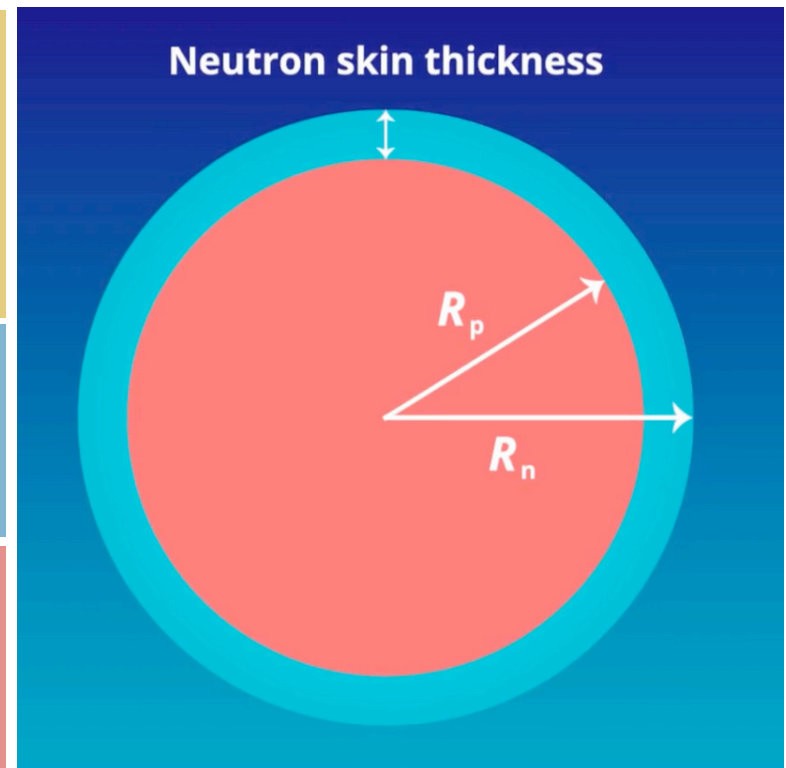
$$\langle r_q^2 \rangle = \int d\vec{r} r^2 \rho_q(\vec{r})$$

**The neutron skin thickness** is defined as the difference between the neutron and proton radii

$$\Delta r_{np} \equiv \langle r_n^2 \rangle^{1/2} - \langle r_p^2 \rangle^{1/2}$$

For **atomic nuclei** with  $N > Z$  we expect a **neutron skin**, **larger** as **larger** the average repulsion (**pressure**) felt by the neutrons

In **first approximation** (within a Local Density Approximation) **the pressure** that a **neutron uniform system feels** at the densities typical for the atomic nucleus





# Neutron skin thickness in atomic nuclei

## Why is it important?

### Constraints the nuclear equation of state (EoS)

The neutron skin is strongly linked to the density dependence of the energy per particle in a uniform system

**[WE WILL DISCUSS THIS IN WHAT FOLLOWS]**

### Connects nuclei to neutron stars

The same physics governing neutron skins also determines properties of neutron stars related to their size and deformability (quadruple polarizability) as well as how exotic are nuclei present in their crusts.

### Tests nuclear structure models

Precise measurements discriminate between competing theoretical models, improving our understanding of nuclear forces.

### Impacts reactions and stability of exotic nuclei

Neutron distribution affects reaction cross sections, decay properties, and predictions for very neutron-rich nuclei (relevant for rare-isotope physics).

### Informs r-process nucleosynthesis

Better constraints on neutron-rich matter improve models of heavy element formation in astrophysical environments.

### Provides electroweak physics tests

Techniques like parity-violating electron scattering probe neutron distributions with minimal strong-interaction uncertainties, offering clean tests of the Standard Model in nuclei.

# Neutron skin thickness in atomic nuclei

How it can be calculated (for any nucleus)?

## DENSITY FUNCTIONAL THEORY (DFT)

### Hohenberg-Kohn theorems

P.Hohenberg, W. Kohn, Phys. Rev. 136, B864 (1964)

→ Assuming a system of **interacting fermions** in a confining **external potential**, there exist a **universal functional  $F[\rho]$**  of the fermion density  $\rho$ :

$$E[\rho] = \langle \Psi | T + V + V_{\text{ext}} | \Psi \rangle = F[\rho] + \int V_{\text{ext}}(r) \rho(r) d\vec{r}$$

→ and it can be shown that

$$\min_{\Psi} \langle \Psi | T + V + V_{\text{ext}} | \Psi \rangle = \min_{\rho} E[\rho]$$

so  $E[\rho]$  has a **minimum** for the **exact ground-state density** where it assumes the **exact energy** as a value.

Nuclei are **self-bound** systems  $\Rightarrow V_{\text{ext}} = 0$

# Kohn-Sham (KS) realization

$$F[\rho] \rightarrow T_{\text{non-int.}}[\rho] + V_{\text{KS}}[\rho]$$

In nuclei no need of external confining potential

For any interacting system, there exists a **local single-particle potential**  $V_{\text{KS}}(\mathbf{r})$ , such that the **exact ground-state density of the interacting system** equals the **ground-state density of the auxiliary non-interacting system**:

$$\rho_{\text{exact}}(\vec{r}) = \rho_{\text{KS}}(\vec{r}) = \sum_{i=1}^A |\phi(\vec{r})|^2$$

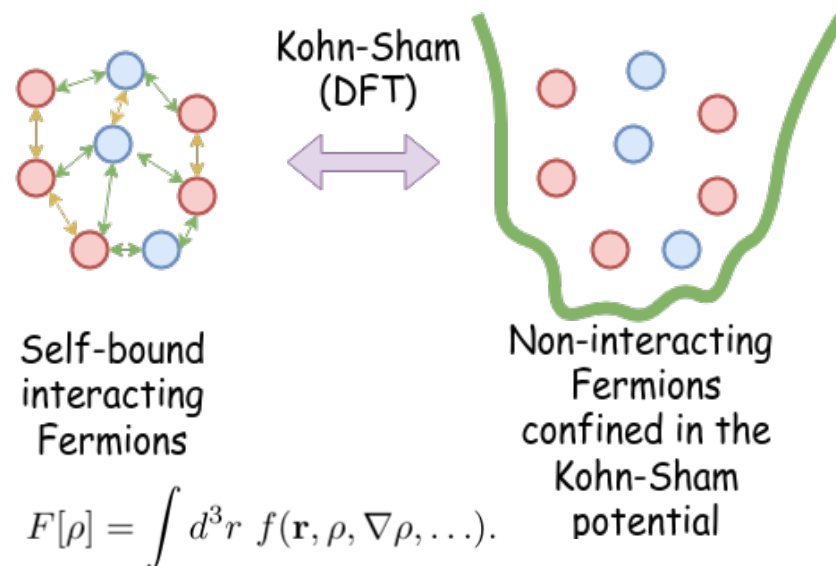
where  $\phi$  are single-particle orbitals and the total wave-function correspond to a Slater determinant. The  $\mathbf{E}[\rho]$  is **unique**

$$E[\rho] = T[\rho] + \int V_{\text{KS}}(\vec{r}) \rho(\vec{r}) d\vec{r}$$

where  $\mathbf{T}[\rho]$  is the **kinetic energy of the non-interacting system** and for which the variational equation

$$0 = \frac{\delta E[\rho]}{\delta \rho} = \frac{\delta T[\rho]}{\delta \rho} + V_{\text{KS}}$$

yields to the **exact ground state density and energy**



# Nuclear DFT (in practice)

## Exactification of Hartree-Fock

Starting from a Hamiltonian  $\mathcal{H}$  or Lagrangian  $\mathcal{L}$  based on zero-range or finite range effective interactions, one derives an EDF by taking the Hartree-Fock or Hartree expectation value

$$\mathcal{H} = \mathcal{T} + \mathcal{V}$$
$$E[\rho, \vec{\nabla} \rho, \tau, \vec{J}, \dots] \equiv \langle HF | \mathcal{H} | HF \rangle$$

**Example:**  
Skyrme  
effective  
interaction  
[zero-range]

$$\begin{aligned} V(\mathbf{r}_1, \mathbf{r}_2) = & t_0 (1 + x_0 P_\sigma) \delta(\mathbf{r}) && \text{central term} \\ & + \frac{1}{2} t_1 (1 + x_1 P_\sigma) \left[ \mathbf{P}'^2 \delta(\mathbf{r}) + \delta(\mathbf{r}) \mathbf{P}^2 \right] \\ & + t_2 (1 + x_2 P_\sigma) \mathbf{P}' \cdot \delta(\mathbf{r}) \mathbf{P} && \text{non-local terms} \\ & + \frac{1}{6} t_3 (1 + x_3 P_\sigma) [\rho(\mathbf{R})]^\sigma \delta(\mathbf{r}) && \text{density-dependent term} \\ & + iW_0 \boldsymbol{\sigma} \cdot \left[ \mathbf{P}' \times \delta(\mathbf{r}) \mathbf{P} \right] && \text{spin-orbit term.} \end{aligned}$$

[  $t_i$ ,  $x_i$  and  $W_0$  parameters to be adjusted to  $B$ ,  $R_{ch}$ , ... ]

With this type of **model** one can calculate the properties of **finite nuclei** and of a **uniform ideal system**

# Nuclear Equation of State (EoS)

What is it? How is defined?

**Definition:** the energy ( $E$ ) per nucleon ( $A=N+Z$ ),  $e \equiv E/A$ , of an **uniform system of neutrons ( $N$ ) and protons ( $Z$ )** as a function of the **neutron ( $\rho_n = N/V$ ) and proton ( $\rho_p = Z/V$ ) densities ...**

→ ... **at zero temperature:** room temperature  $10^2\text{K} \rightarrow 10^{-8}$  MeV while “cold” neutron stars are at about  $10^{10}\text{K} \rightarrow 1$  MeV. **Separation energy** in stable nuclei (equivalent to ionization energy in atoms) is of **several MeV**.

→ ... **unpolarized:** energy favors **couples** of neutrons and protons **occupying the same state but with opposite spins** (equivalent to electrons in atoms)

→ ... **isospin symmetric:** neutron-neutron, proton-proton and neutron-proton **nuclear interaction** are very **similar** among them. **Masses** of neutrons and protons are almost **degenerate**. Hence neutrons and protons can be thought as **two states** of the **same particle** with different isobaric spin or **isospin** (in analogy with spin): the **nucleon**.

→ ... and **no Coulomb:** **idealized uniform system** (focus on strong interaction, much stronger than Coulomb at nuclear scale  $\sim 10^{15}$  fm). Real systems are finite and electrically neutral so no divergences in adding Coulomb.

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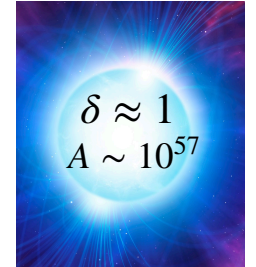
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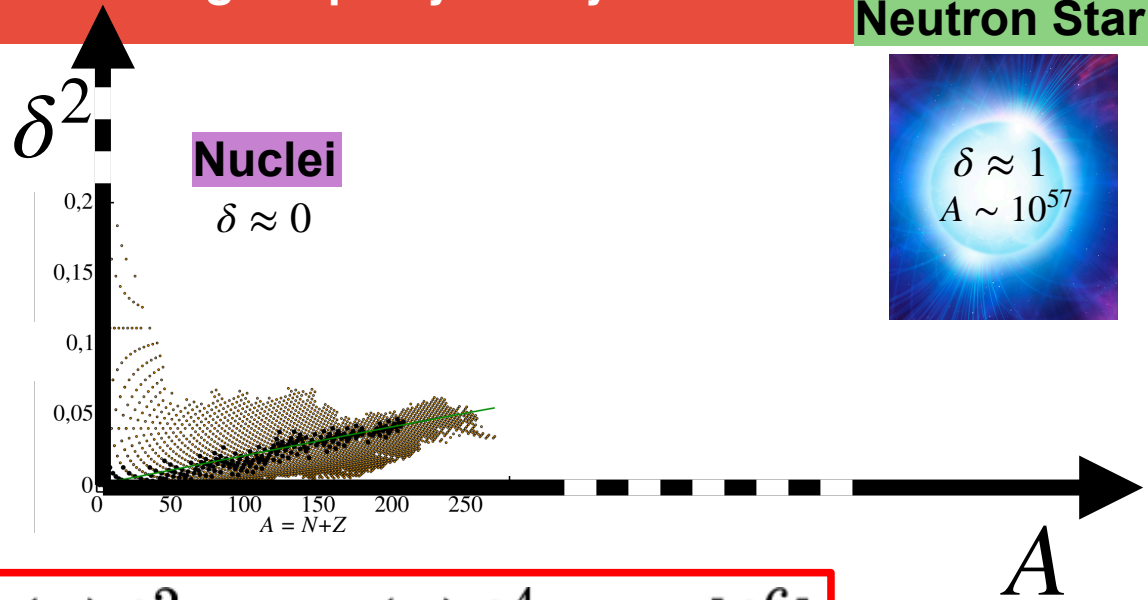
# Nuclear Equation of State (EoS)

Unpolarized, uniform nuclear matter at  $T=0$  assuming isospin symmetry

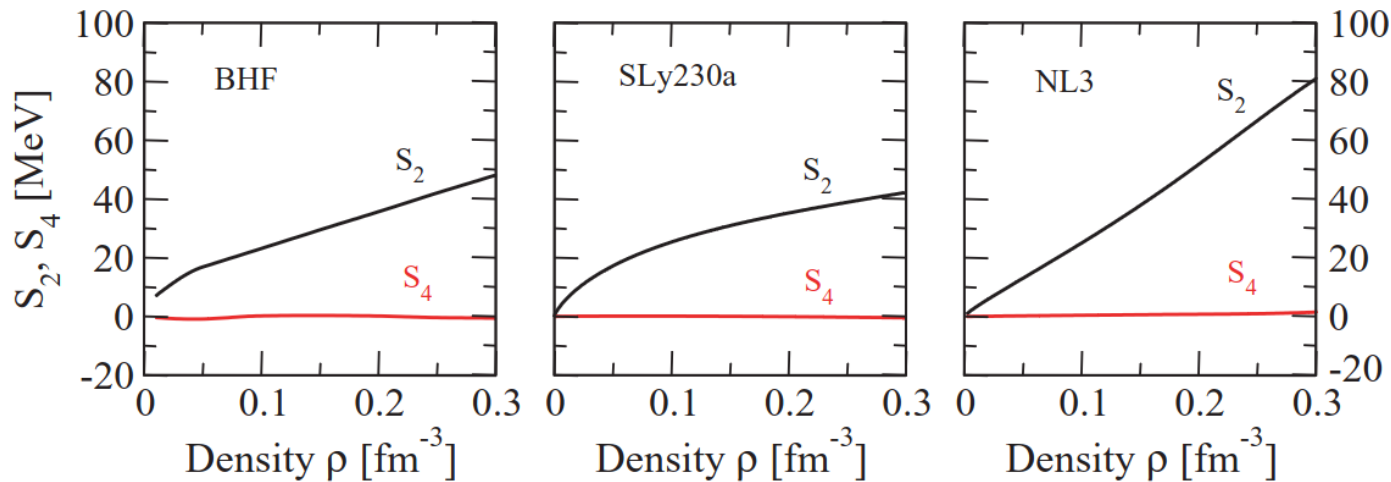
Neutron Star



It is convenient to write the energy per nucleon ( $e$ ) as a function of the total density  $\rho = \rho_n + \rho_p$  and the relative difference  $\delta = (\rho_n - \rho_p) / \rho$  for  $\delta \rightarrow 0$ :



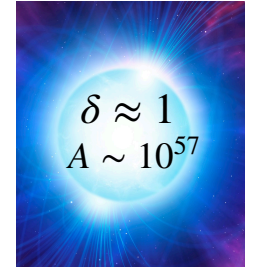
$$e(\rho, \delta) = e(\rho, 0) + S_2(\rho)\delta^2 + S_4(\rho)\delta^4 + \mathcal{O}[\delta^6]$$



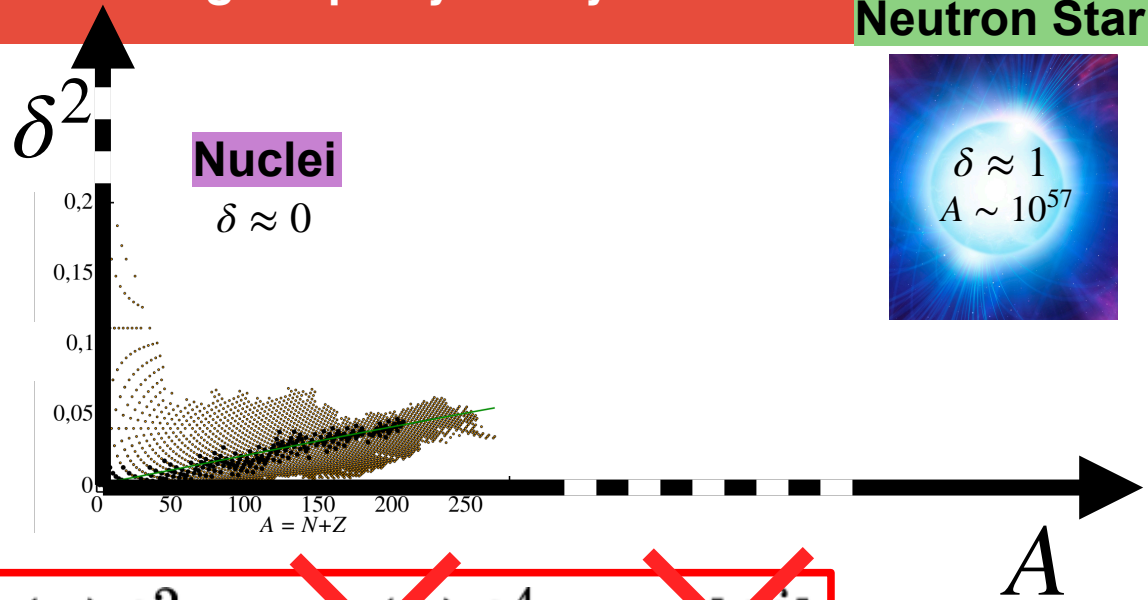
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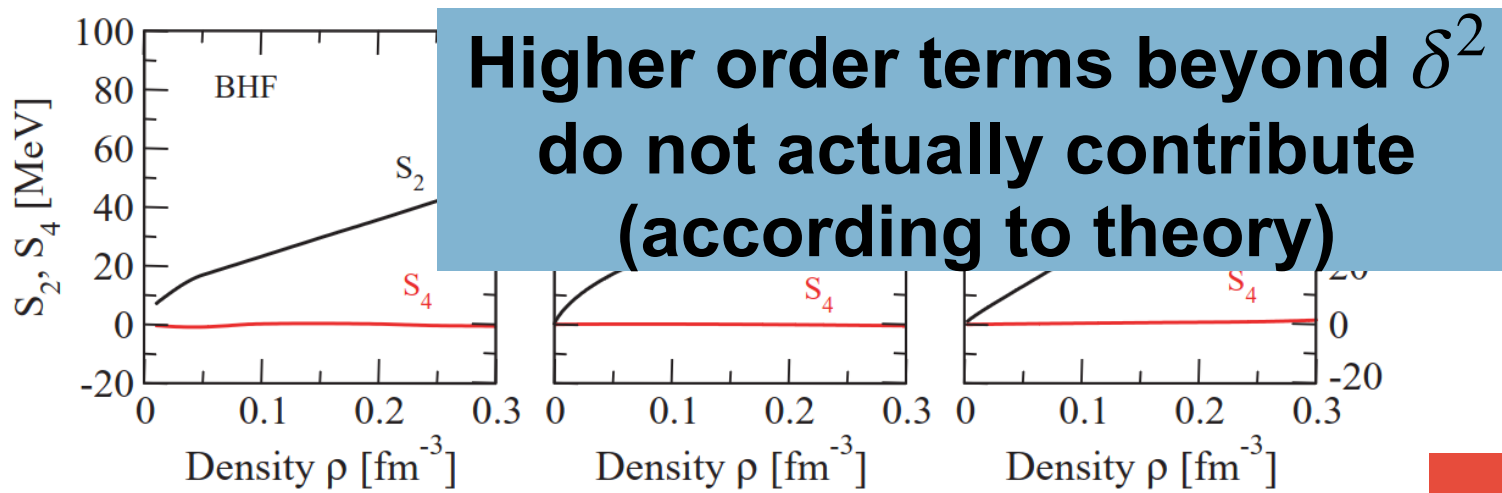
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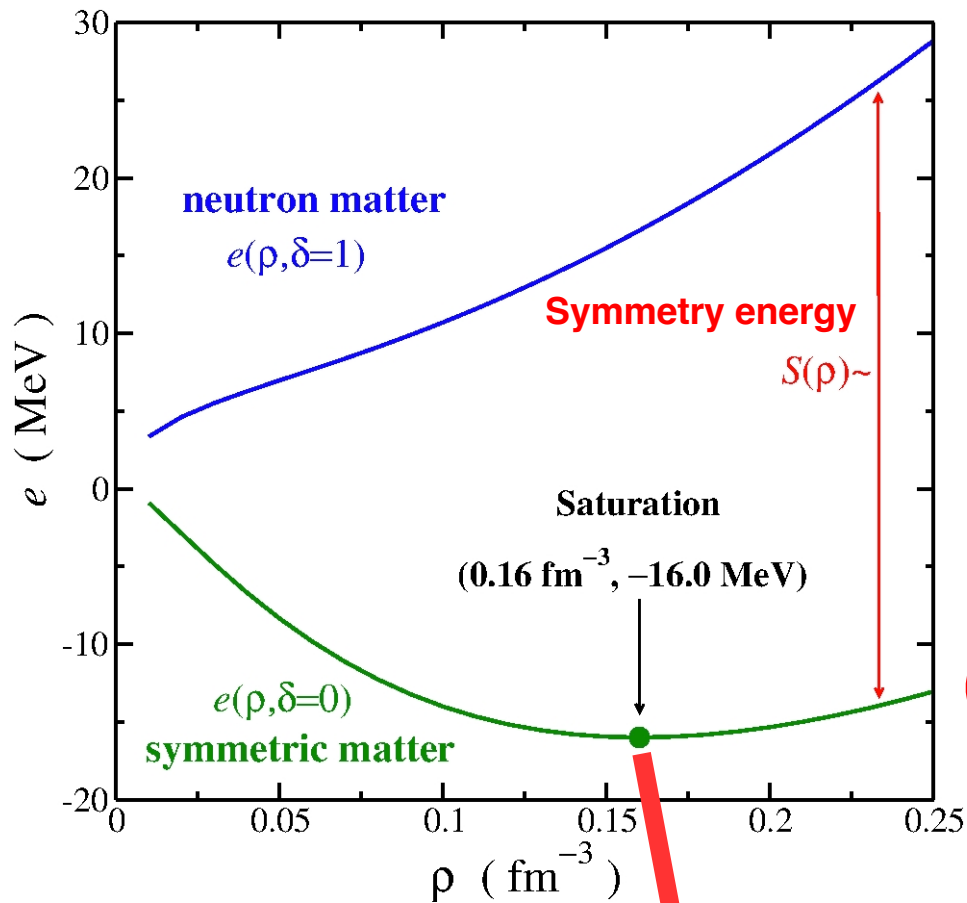
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# Nuclear Equation of State (EoS)

Unpolarized, uniform nuclear matter at T=0 assuming isospin symmetry

$$e(\rho, \delta) = e(\rho, 0) + S_2(\rho)\delta^2$$



It is customary to **Taylor expand**  $e(\rho, 0)$  and  $S(\rho)$  around **nuclear saturation density**  $\rho_0 \sim 0.16 \text{ fm}^{-3}$

$$e(\rho, 0) = e(\rho_0, 0) + \frac{1}{2}K_0x^2 + \mathcal{O}[\rho^3] \text{ where } x = \frac{\rho - \rho_0}{3\rho_0}$$

$$S(\rho) = J + Lx + \frac{1}{2}K_{\text{sym}}x^2 + \mathcal{O}[\rho^3, \delta^4]$$

**K** → how **compressible** is matter @  $\rho_0$

**J** → **penalty energy** for systems with  $\rho_n \neq \rho_p$  @  $\rho_n + \rho_p = \rho_0$

**L** → **pressure** for systems with  $\rho_n \neq \rho_p$  @  $\rho_0$

$$P = - \left. \frac{\partial E}{\partial V} \right|_{A=\text{const.}} = \frac{1}{3} \rho \delta^2 L(\rho)$$

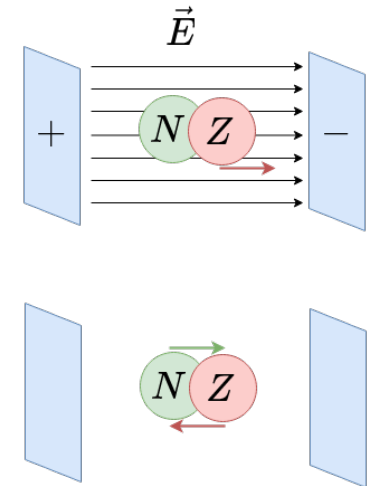
# Electric Dipole Polarizability: introduction

The **electric dipole polarizability** measures the **tendency** of the nuclear **charge distribution** to be **distorted**

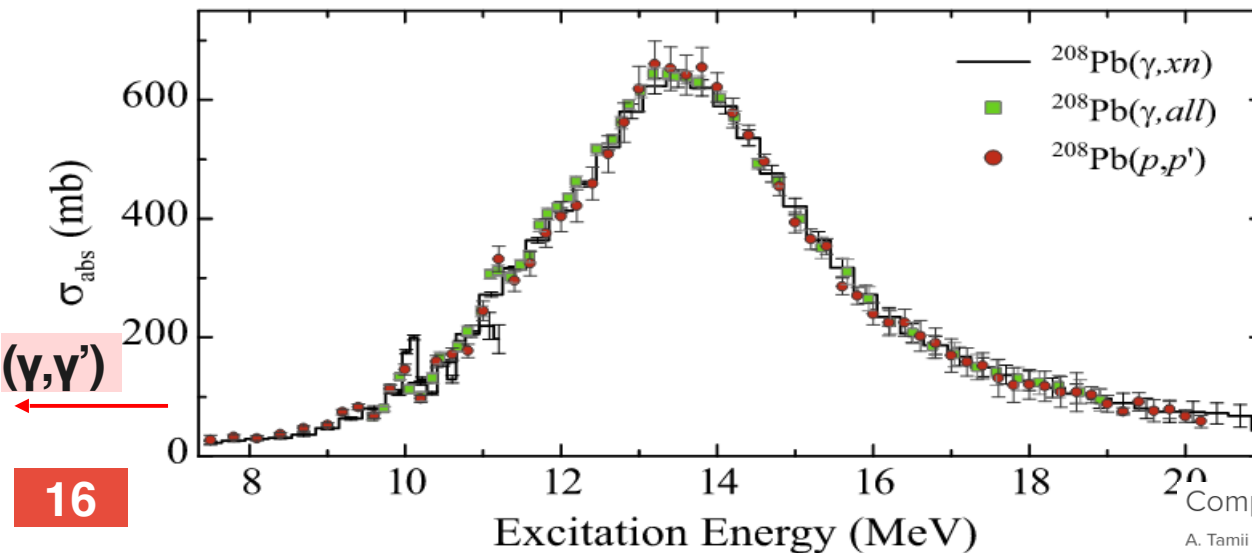
$$\alpha = \frac{\text{electric dipole moment}}{\text{external electric field applied}}$$

Microscopically, it relates with the **photo-absorption cross-section**

$$\alpha_D = \frac{\hbar c}{2\pi^2 e^2} \int \frac{\sigma_{\text{abs}}}{\omega^2} d\omega,$$



Measured using **polarized** proton scattering at **very forward angles** (dominated by **E1** and **M1 well separated**)



$$\alpha_D(^{208}\text{Pb}) = 19.6 \pm 0.6 \text{ fm}^3$$

$$\Delta r_{\text{np}} = 0.156^{+0.025}_{-0.021} \text{ fm}$$

# Electric Dipole Polarizability: theory

Theoretically, the total photo-absorption cross section, can be written as

$$\sigma_{\gamma\text{-abs}} = 4\pi^2\alpha \sum_{\nu} (E_{\nu} - E_0) |\langle \nu | F_{\text{dipole}} | 0 \rangle|^2$$

Dipole operator  
subtract CM motion

And, thus,

$$\alpha_D = 2 \sum_{\nu \neq 0} \frac{|\langle \nu | F_{\text{dipole}} | 0 \rangle|^2}{E_{\nu} - E_0} \equiv 2m_{-1} \quad \frac{eN}{A} \sum_{i=1}^Z r_i Y_{1M}(\hat{r}_i) - \frac{eZ}{A} \sum_{i=1}^N r_i Y_{1M}(\hat{r}_i)$$

Considering the G.S. perturbed by an external field  $\lambda F$  (with  $\lambda \rightarrow 0$ ):

$\langle \mathcal{H} \rangle = \langle \mathcal{H}_0 + \lambda F_{\text{dipole}} \rangle$ ; The variation in the expectation energy can be written as:

$$\delta \langle \mathcal{H} \rangle = \lambda^2 \sum_{\nu \neq 0} \frac{|\langle \nu | F | 0 \rangle|^2}{E_{\nu} - E_0} + \mathcal{O}(\lambda^3) = \lambda^2 m_{-1} + \mathcal{O}(\lambda^3)$$

$$m_{-1} = \frac{1}{2} \frac{\partial^2 \langle \mathcal{H} \rangle}{\partial \lambda^2} \Big|_{\lambda=0}$$

Dielectric  
Theorem

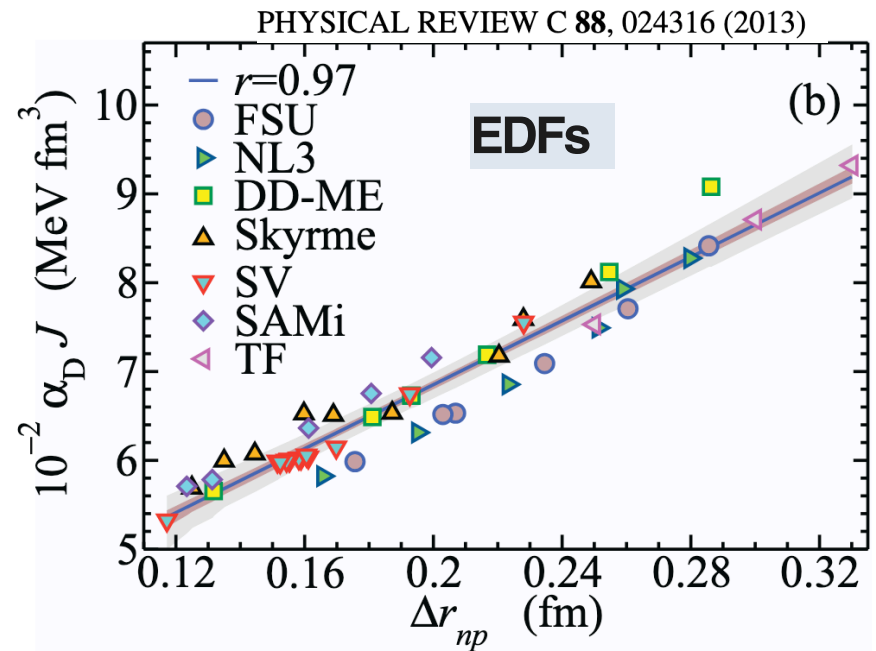
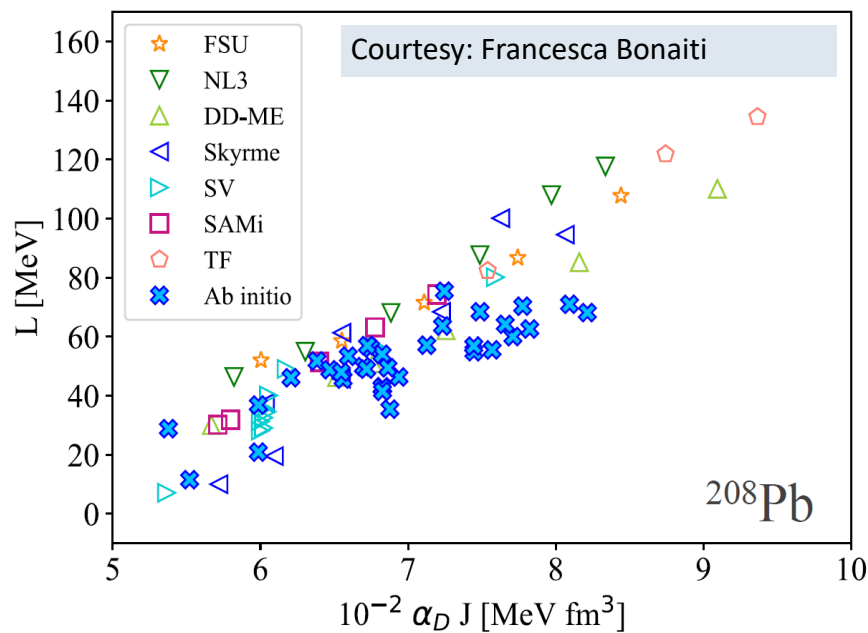
# Electric Dipole Polarizability: simple model & correlations

Applying the **dielectric theorem** to the **Droplet Model** Hamiltonian (first Migdal and latter on Meyer et al. NPA385, 269) one can find

$$\alpha_D \approx \frac{A \langle r^2 \rangle}{12J} \left[ 1 + \frac{5 \Delta r_{np} + \sqrt{\frac{3}{5} \frac{e^2 Z}{70J}} - \Delta r_{np}^{\text{surface}}}{\langle r^2 \rangle^{1/2} (I - I_C)} \right]$$

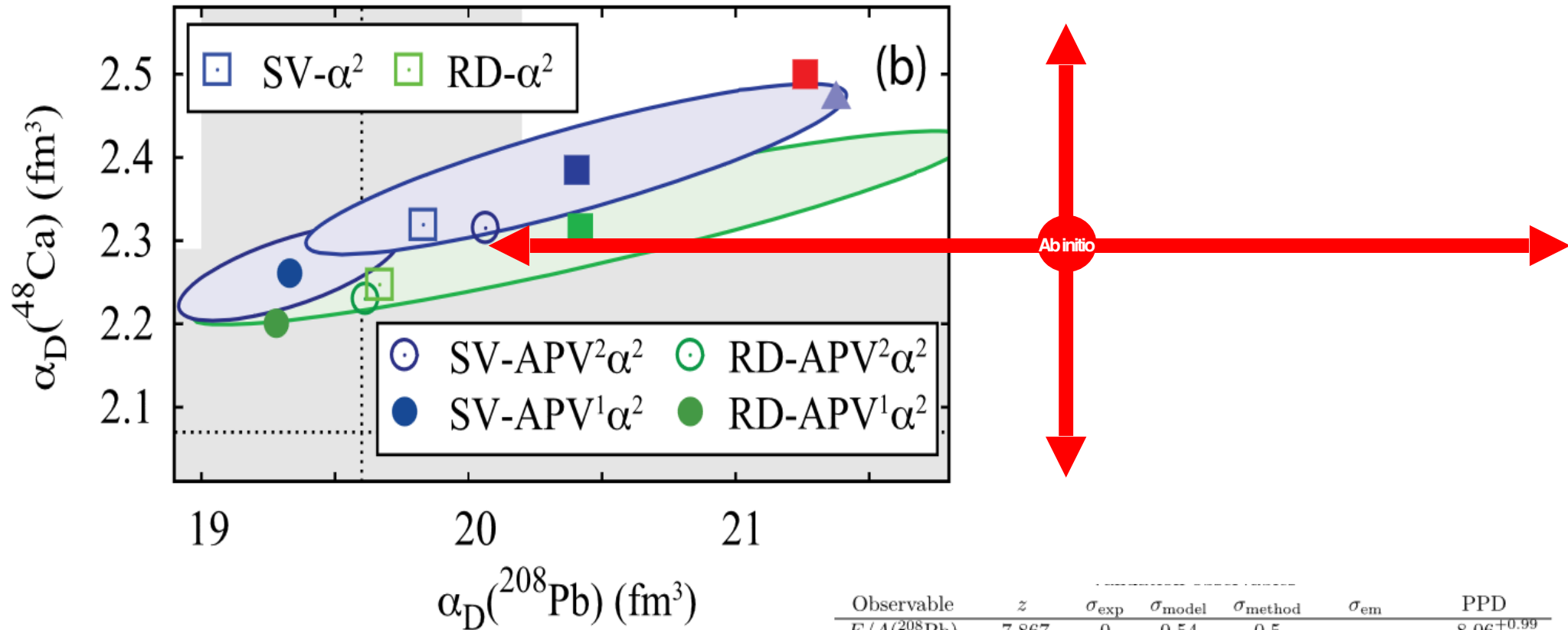
$J = e(\text{PNM}) - e(\text{SNM}) \rightarrow$  Symmetry energy at  $\rho_0$   
 $\Delta r_{np} \equiv \langle r_n^2 \rangle^{1/2} - \langle r_p^2 \rangle^{1/2} \rightarrow$  Neutron skin thickness

## Using microscopic calculations (Energy Density Functionals & Ab initio)



$$\alpha_D J = (301 \pm 32) + (1922 \pm 73) \Delta r_{np}$$

# Electric Dipole Polarizability in $^{48}\text{Ca}$ and $^{208}\text{Pb}$



Combined Theoretical Analysis of the Parity-Violating Asymmetry for  $^{48}\text{Ca}$  and  $^{208}\text{Pb}$

Paul-Gerhard Reinhard, Xavier Roca-Maza, and Witold Nazarewicz  
Phys. Rev. Lett. **129**, 232501 – Published 2 December 2022

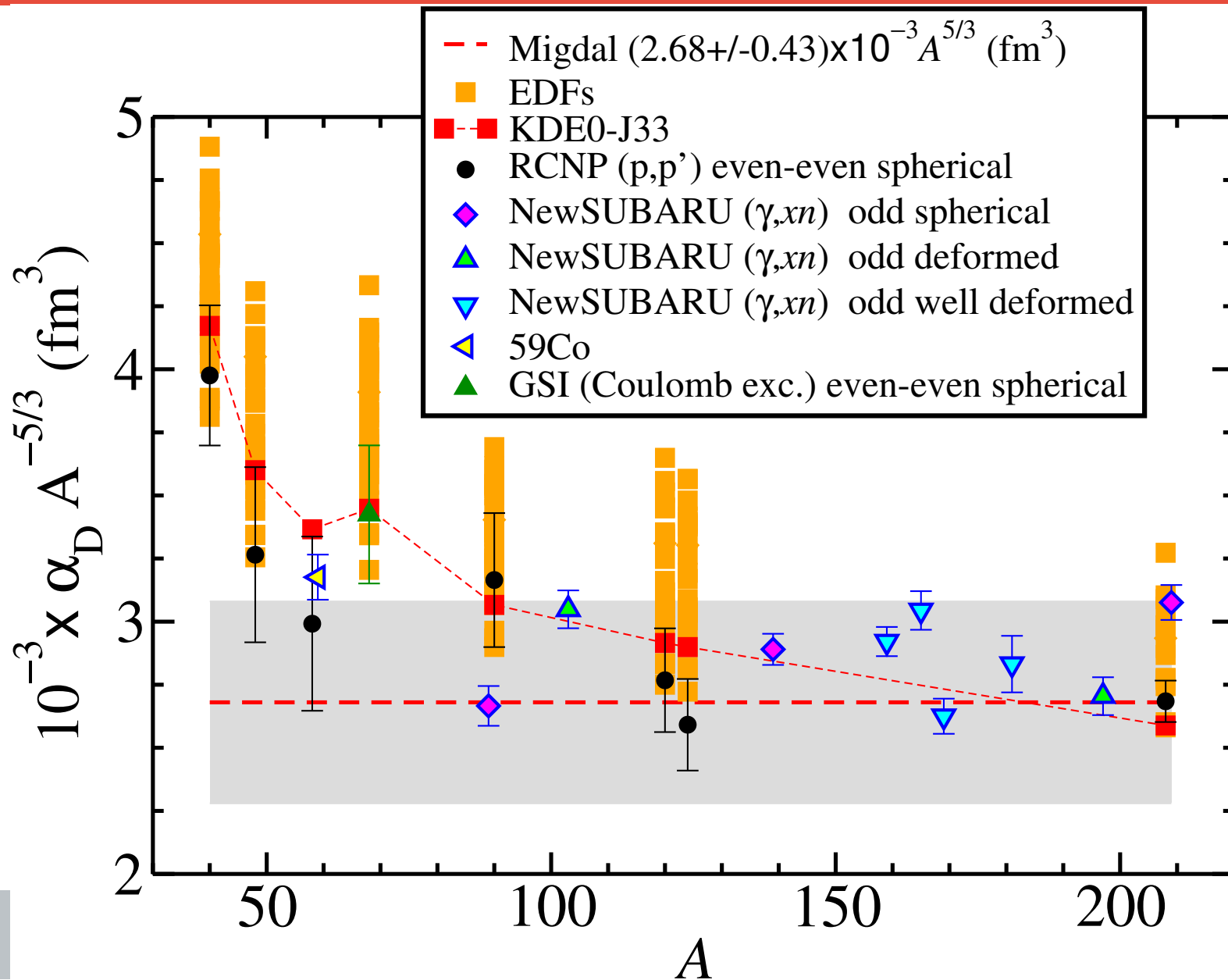
| Observable                  | $z$    | $\sigma_{\text{exp}}$ | $\sigma_{\text{model}}$ | $\sigma_{\text{method}}$ | $\sigma_{\text{em}}$ | PPD                     |
|-----------------------------|--------|-----------------------|-------------------------|--------------------------|----------------------|-------------------------|
| $E/A(^{208}\text{Pb})$      | -7.867 | 0                     | 0.54                    | 0.5                      | —                    | $-8.06^{+0.99}_{-0.88}$ |
| $R_p(^{208}\text{Pb})$      | 5.45   | 0                     | 0.17                    | 0.05                     | —                    | $5.43^{+0.21}_{-0.23}$  |
| $\alpha_D(^{48}\text{Ca})$  | 2.07   | 0.22                  | 0.06                    | 0.1                      | —                    | $2.30^{+0.31}_{-0.26}$  |
| $\alpha_D(^{208}\text{Pb})$ | 20.1   | 0.6                   | 0.59                    | 0.8                      | —                    | $22.6^{+2.1}_{-1.8}$    |

**Ab initio predictions link the neutron skin of  $^{208}\text{Pb}$  to nuclear forces**

Baishan Hu, Weiguang Jiang, Takayuki Miyagi, Zhonghao Sun, Andreas Ekström, Christian Forssén, Gaute Hagen, Jason D. Holt, Thomas Papenbrock, S. Ragnar Stroberg & Ian Vernon

Nature Physics **18**, 1196–1200 (2022) | [Cite this article](#)

# Electric Dipole Polarizability: How nuclear models perform for other nuclei?



# Parity Violating Asymmetry: introduction

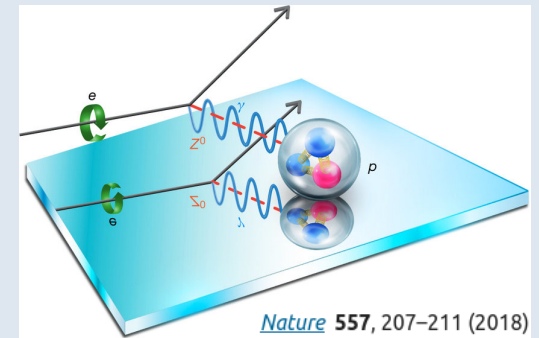
**Elastic electron scattering** by nuclei:

→ **Parity conserving**

Exchange of  $\gamma$  that essentially couples to protons  $Q_p^C=1; Q_n^C=0$

→ **Parity violating**

Exchange  $Z_0$  that essentially couples to neutrons  $Q_p^W=0.07; Q_n^W=-0.99$



**Ultra-relativistic electrons** ( $m_e \rightarrow 0$ ) with spin **aligned** (+) or **anti-aligned** (-) to the beam line, in approaching the nucleus:

$$\left[ \vec{\alpha} \cdot \vec{p} + V_{\text{Coulomb}}(\vec{r}) \pm V_{\text{Weak}}(\vec{r}) \right] \Psi_{\pm} = E \Psi_{\pm}$$

$$\sigma \approx \left| \begin{array}{c} \text{Diagram 1: } e^- \text{ and } 208\text{Pb} \text{ connected by a } \gamma \text{ line} \\ + \\ \text{Diagram 2: } e^- \text{ and } 208\text{Pb} \text{ connected by a } Z^0 \text{ line} \end{array} \right|^2$$

One can get **advantage** from this **interference pattern** between **electromag.** and **weak** interact.

$$A_{PV} = \frac{\frac{d\sigma_+}{d\Omega} - \frac{d\sigma_-}{d\Omega}}{\frac{d\sigma_+}{d\Omega} + \frac{d\sigma_-}{d\Omega}} \approx \frac{\text{Weak}}{\text{Coulomb}}$$

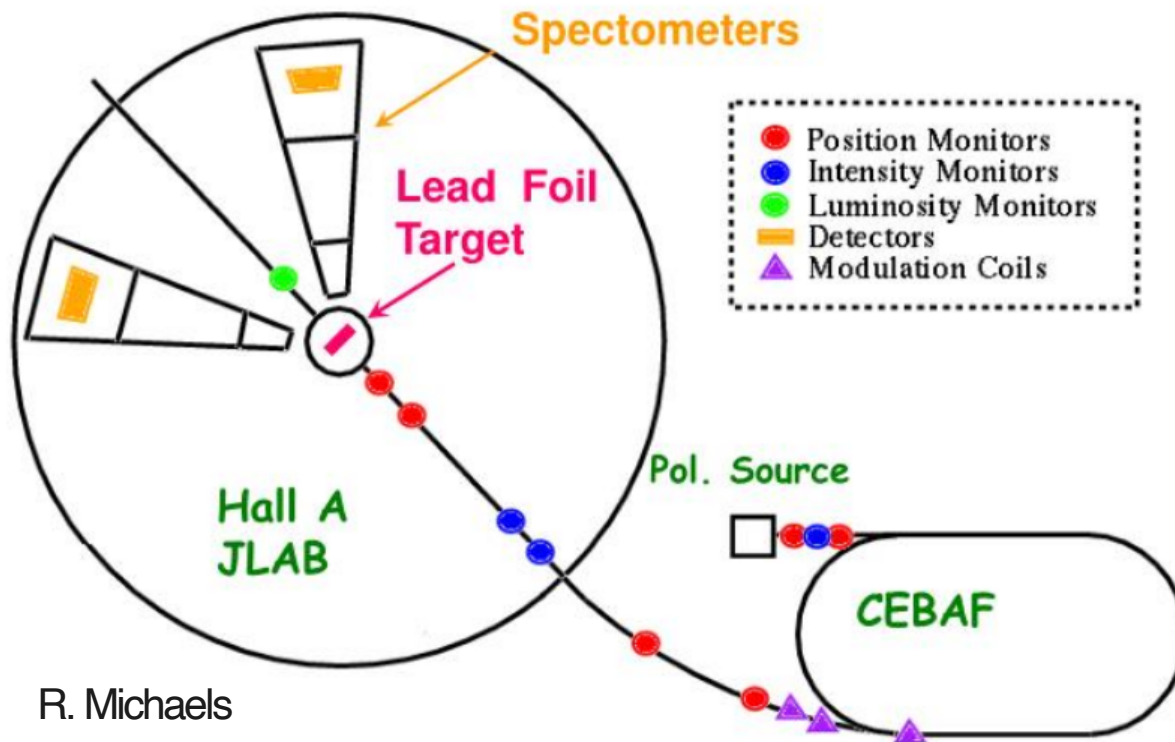
# Parity Violating Asymmetry: experiment

## Lead and Calcium

## Radius Experiments

## PREx & CREx @ JLab

$$A_{PV} \sim \frac{G_F}{4\sqrt{2}\pi\alpha} Q^2 \sim 10^{-4} (Q[\text{GeV}])^2$$



R. Michaels

→ Spectrometers  $\sim 5^\circ$

→  $e^-$  beam  $E \sim \text{GeV}$

$$Q_{\text{PREx}}^2 = 0.00616 \text{ GeV}^2$$

$$Q_{\text{CREx}}^2 = 0.0297 \text{ GeV}^2$$

→ Demanding experiment:

For  $10^6$  electrons or more only **one** of them **interact weakly** with the nucleus

**New experiment on  $^{208}\text{Pb}$  @ Mainz !!! (MREX)**

# Parity Violating Asymmetry: theory

To solve the **Dirac equation**

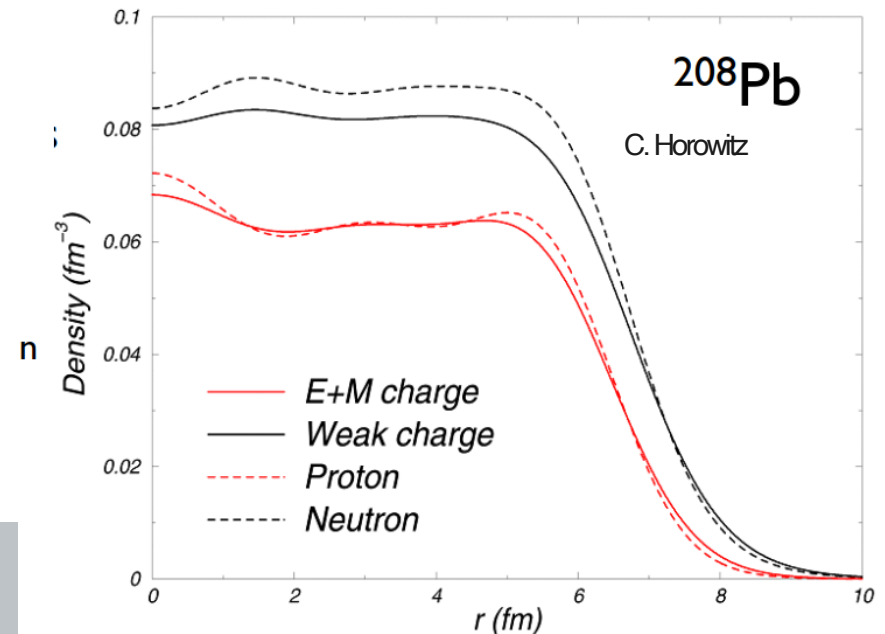
$$\left[ \vec{\alpha} \cdot \vec{p} + V_{\text{Coulomb}}(\vec{r}) \pm V_{\text{Weak}}(\vec{r}) \right] \Psi_{\pm} = E \Psi_{\pm}$$

The main inputs are the **electromagnetic**  $\rho_{\text{ch}}$  and **weak charge**  $\rho_{\text{W}}$  distributions

$$V_{\text{Coulomb}}(\vec{r}) = Z\alpha \int \frac{\rho_{\text{ch}}(\vec{r}')}{|\vec{r} - \vec{r}'|} d\vec{r}'$$

$$V_{\text{Weak}}(\vec{r}) = \frac{G_F}{2\sqrt{2}} \rho_{\text{W}}(\vec{r})$$

We need to know the **neutron and proton distributions** in nuclei and the **electromagnetic and weak charge form factors** of the **neutron** and the **proton**



# Parity Violating Asimmetry: simple model & correlations

## The parity violating asymmetry within the PWBA:

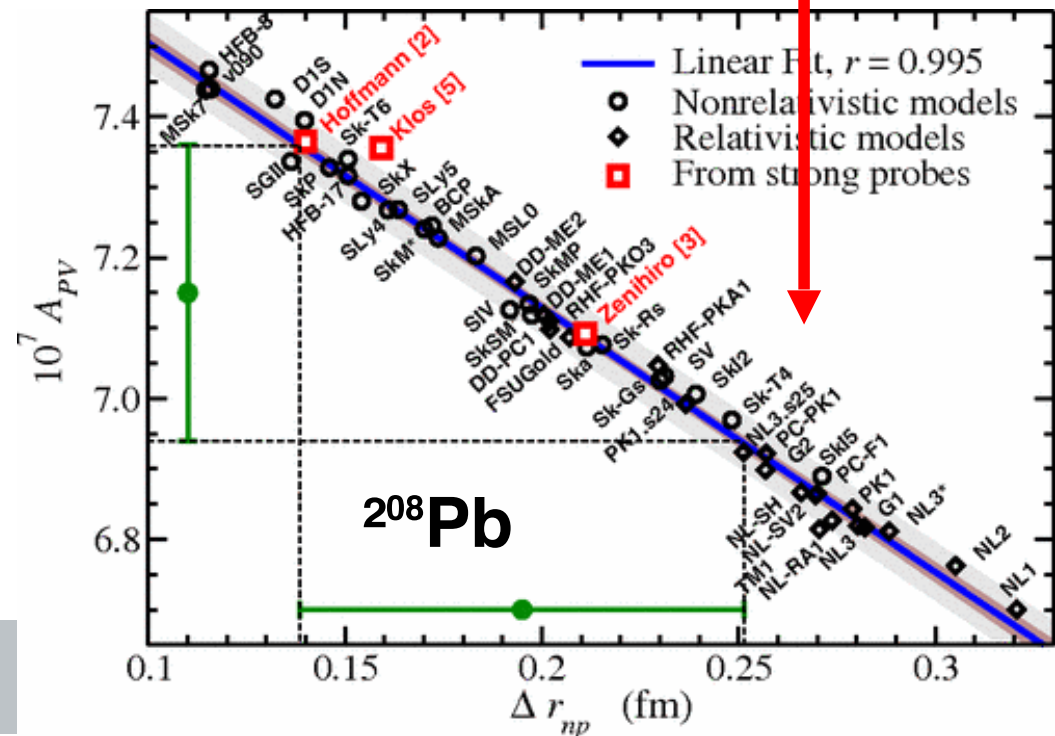
$$A_{PV}^{PWBA} = \frac{G_F Q^2}{4\sqrt{2}\pi\alpha} \left[ \underbrace{4 \sin^2 \theta_W}_{\approx 1} + \frac{F_n(Q) - F_p(Q)}{F_p(Q)} \right]$$

$$\xrightarrow{Q \rightarrow 0} \frac{G_F Q^2}{4\sqrt{2}\pi\alpha} \left[ 1 - \frac{Q^2 \langle r_p^2 \rangle}{3} \frac{\Delta r_{np}}{\langle r_p^2 \rangle^{1/2}} \right]$$

30% due to Coulomb distortions (PREX) does not blur the main physics

## The parity violating asymmetry within the DWBA:

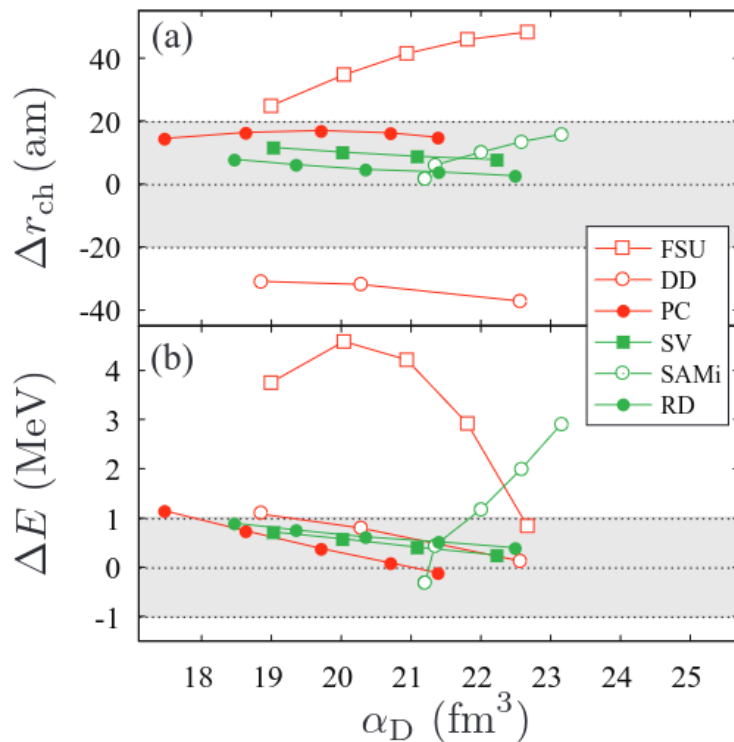
- ◆ → Energy Density Functionals
- → Expeirmental  $\rho_n$  form hadronic probes +  $\rho_p$  from  $e^-$  scattering



# Parity Violating Asymmetry: measurements vs. theory

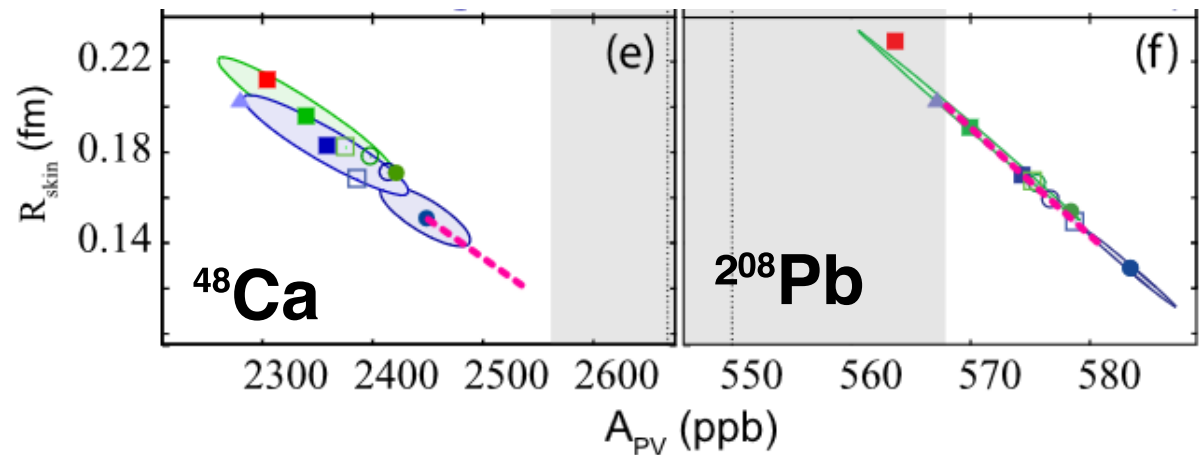
Selecting those models (we consider) well calibrated to masses and radii

Theoretical (EDFs and *ab initio*) and experimental  $1\sigma$  errors overlap in  $^{208}\text{Pb}$  but not in  $^{48}\text{Ca}$  → **No simultaneous description**



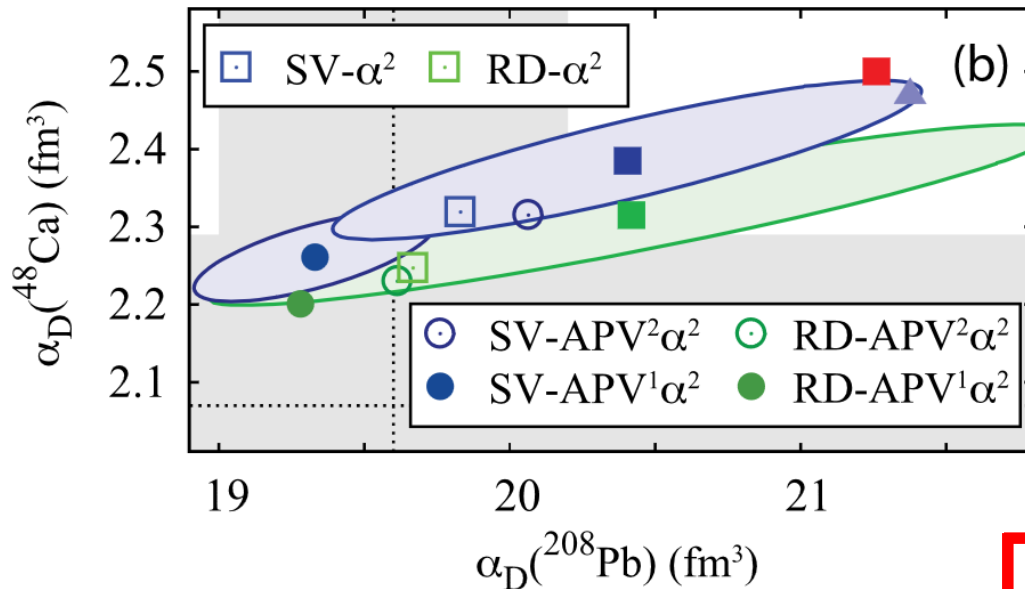
Fitting protocol also including  $A_{PV}$  and  $\alpha_D$

- SV-min
- SV-APV $^2\alpha^2$
- SV-APV $^1\alpha^2$
- SV- $\alpha^2$
- ▲ SV-min\*
- RD-min
- RD-APV $^2\alpha^2$
- RD-APV $^1\alpha^2$
- RD- $\alpha^2$
- PC-min



# Summary: model performance

$A_{PV}$  (sensitive to  $\Delta r_{np}$ ) and  $\alpha_D$  (sensitive to  $J$  and  $\Delta r_{np}$ ) in  $^{48}\text{Ca}$  and  $^{208}\text{Pb}$



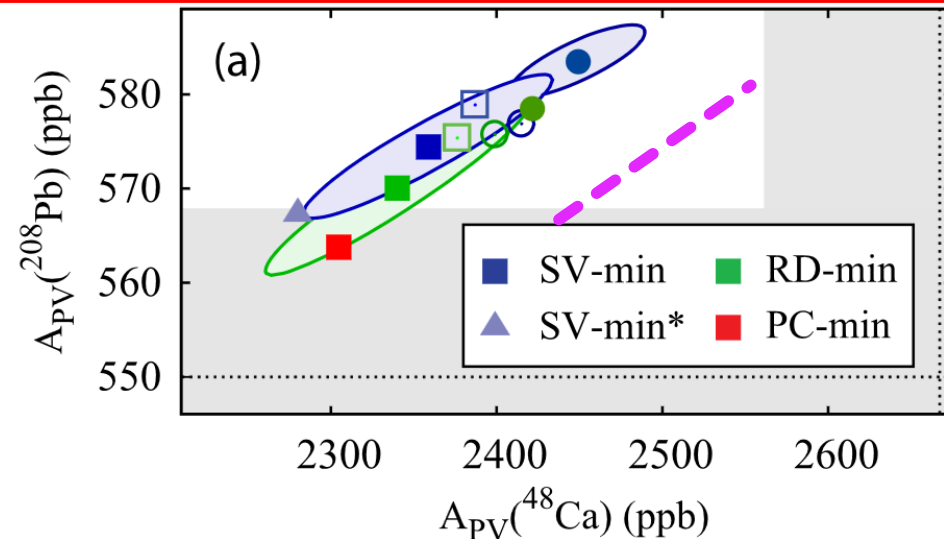
**Simultaneous description of dipole polarizabilities** → point to a **good understanding of symmetry energy (J) and neutron skins ( $\Delta r_{np}$ )**

**Ab-initio (B. Hu) Nature Physics (2022)**

$$\alpha_D(^{48}\text{Ca}) \quad 2.30^{+0.31}_{-0.26}$$

$$\alpha_D(^{208}\text{Pb}) \quad 22.6^{+2.1}_{-1.8}$$

**No simultaneous description of parity violating asymmetries** (ground state observable) → point to a **deficient understanding of neutron skins**



**Magenta dashed lines** from extrapolated  $\Delta r_{np}$  given in G. Hagen et al. Nature Physics 12, 186–190 (2016) and H. Bu et al. Nature Physics (2022)

# $A_{PV}$ theoretical corrections?

**In all analysis:** Coulomb potential has been considered at tree level while  $Q_{Weak}$  has been corrected at one-loop level (interference of the  $\gamma$  and  $Z_0$  exchange):

$$[-i\vec{\alpha} \cdot \vec{\nabla} + V_{Coul} \pm V_{Weak}] \Psi_{\pm} = E \Psi_{\pm}$$

**First order QED includes:** vacuum polarization (VP), self-energy and  $e^-$  vertex correction (V-SE) [Jakubassa-Amundsen (2024) JPG: Nucl. Part. Phys. 51 035105]

$$[-i\vec{\alpha} \cdot \vec{\nabla} + V_{Coul} \pm V_{Weak} + V_{VP} + V_{V-SE}] \Psi_{\pm} = E \Psi_{\pm}$$

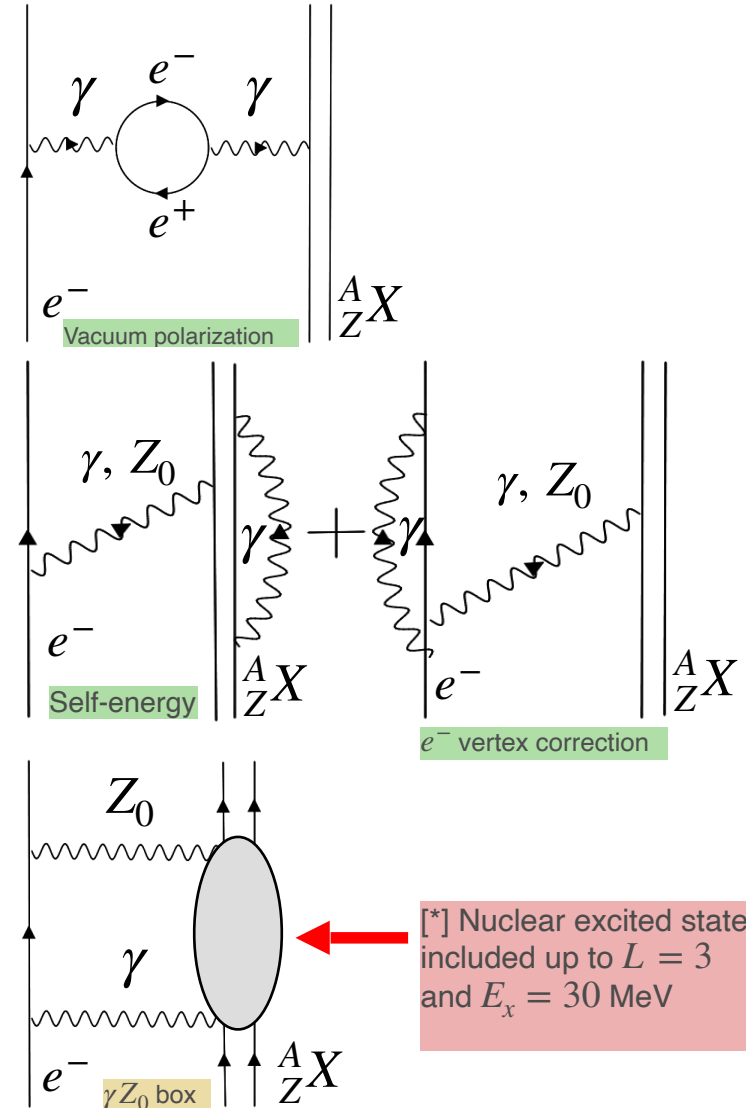
**Dispersion corrections:** interference of the  $\gamma$  and  $Z_0$  exchange have been shown to be relevant for estimating the  $Q_{weak}$ .

On  $A_{PV}$  @ PREx kin., corrections **small & cancel**

- Box diagram  $-0.1\%$
- QED  $0.09\%$

[Jakubassa-Amundsen and XRM, arXiv:2507.15380]

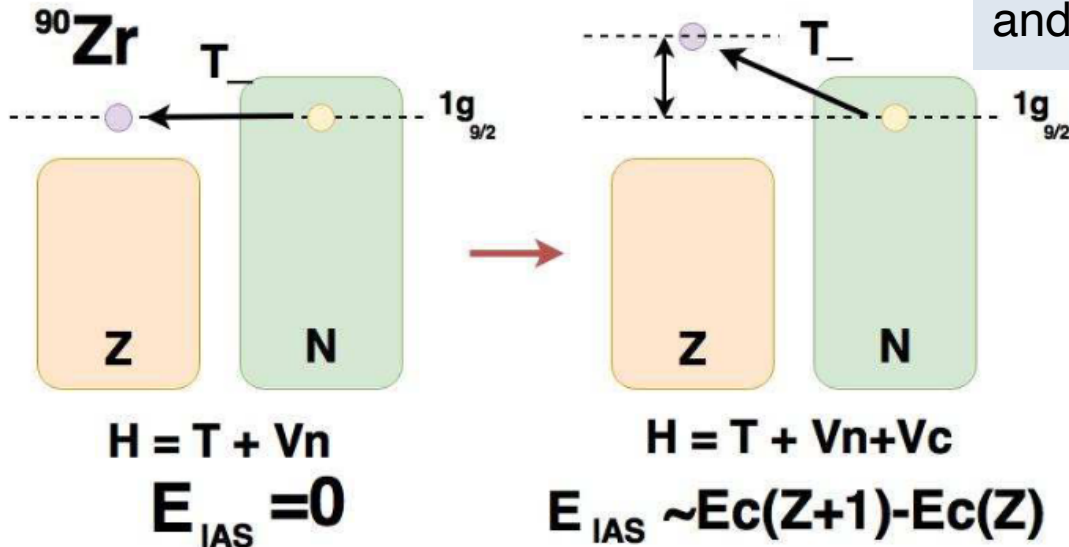
[XRM and Jakubassa-Amundsen PRL 134, 192501 (2025)]



# Isobaric Analog State: Example + theory

$$F = T_{\pm} = \sum_i^A t_{\pm}(i)$$

## Charge-exchange experiments



Isospin algebra analogous to spin algebra  $s \rightarrow t$  and  $\tau \rightarrow \sigma$  (Pauli matrices with  $t = \tau/2$ )

$$t_- |n\rangle = \frac{1}{2} |p\rangle$$

$$t_+ |p\rangle = -\frac{1}{2} |n\rangle$$

$$T_+^\dagger = T_- \quad T_-^\dagger = T_+$$

$$[T_z, T_{\pm}] = \pm T_{\pm}$$

$$[T_+, T_-] = 2T_z$$

→ non-energy weighted sum rule:

$$\begin{aligned}
 m_0^- - m_0^+ &= \langle 0 | T_+ T_- | 0 \rangle - \langle 0 | T_- T_+ | 0 \rangle \\
 &= \langle 0 | [T_+, T_-] | 0 \rangle = \langle 0 | 2T_z | 0 \rangle \\
 &= N - Z
 \end{aligned}$$

Note: If not isospin-mixing it would be zero!!

→ energy weighted sum rule:

$$m_1 = \sum_{\nu} (E_{\nu} - E_0) |\langle \nu | F | 0 \rangle|^2 = \langle 0 | T_+ [\mathcal{H}, T_-] | 0 \rangle$$

$[\mathcal{H}, T_-] \neq 0$  only if  $\mathcal{H}$  contains terms that **breaks isospin invariance**

# Isobaric Analog State: Theory

$$F = T_{\pm} = \sum_i^A t_{\pm}(i)$$

→ Hence, the **centroid energy**  $m_1/m_0$ :

$$E_{\text{IAS}} = \frac{\langle 0 | T_+ [\mathcal{H}, T_-] | 0 \rangle}{\langle 0 | T_+ T_- | 0 \rangle} = \frac{1}{N - Z} \langle 0 | T_+ [\mathcal{H}, T_-] | 0 \rangle$$

→ Assuming a simple model: **independent particle** model with only **Coulomb breaking isospin symmetry** (neglect exchange effects)

$$E_{\text{IAS}}^{\text{C,direct}} = \frac{1}{N - Z} \int [\rho_n(\vec{r}) - \rho_p(\vec{r})] U_{\text{C}}^{\text{direct}}(\vec{r}) d\vec{r},$$

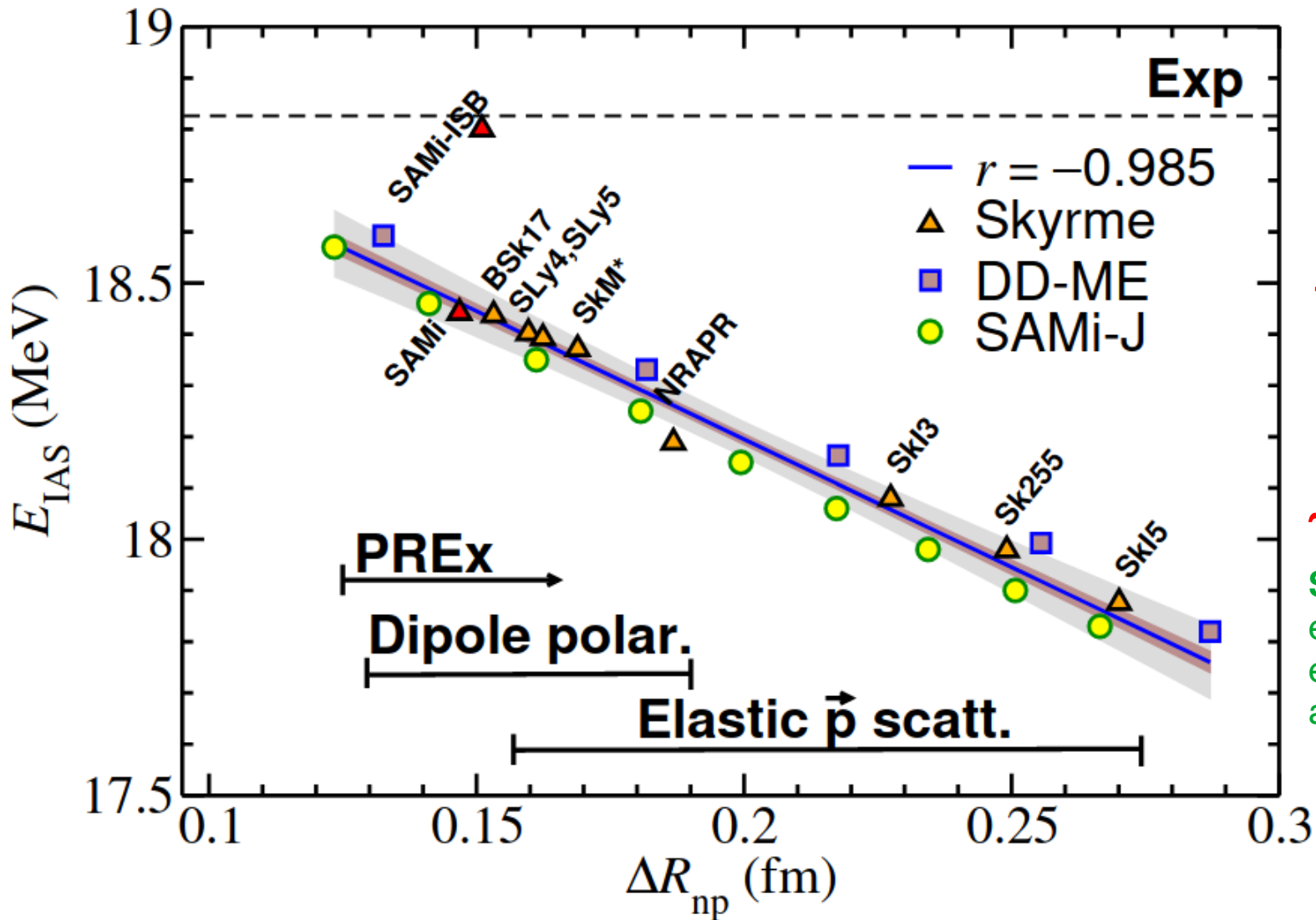
$$U_{\text{C}}^{\text{direct}}(\vec{r}) = \int \frac{e^2}{|\vec{r}_1 - \vec{r}|} \rho_{\text{ch}}(\vec{r}_1) d\vec{r}_1$$

→ Assuming sharp sphere to describe  $\rho_n$  and  $\rho_p$  and  $\rho_{\text{ch}} = \rho_p$

$$U_{\text{C}}^{\text{direct}}(\vec{r}) = \begin{cases} \frac{Ze^2}{2R_p} \left( 3 - \frac{r^2}{R_p^2} \right) & \text{for } r < R_p \\ \frac{Ze^2}{r} & \text{for } r > R_p \end{cases}$$

$$\begin{aligned} E_{\text{IAS}} &\approx E_{\text{IAS}}^{\text{C,direct}} \\ &\approx \frac{6Ze^2}{5R_p} \left( 1 - \frac{1}{2} \frac{N}{N-Z} \frac{R_n - R_p}{R_p} \right) \\ &\approx \frac{6}{5} \frac{Ze^2}{r_0 A^{1/3}} \left( 1 - \sqrt{\frac{5}{12}} \frac{N}{N-Z} \frac{\Delta R_{\text{np}}}{r_0 A^{1/3}} \right) \end{aligned}$$

# Isobaric Analog State: Theory+experiment



Exp errors in IAS ~  
tens of keV

(or smaller)

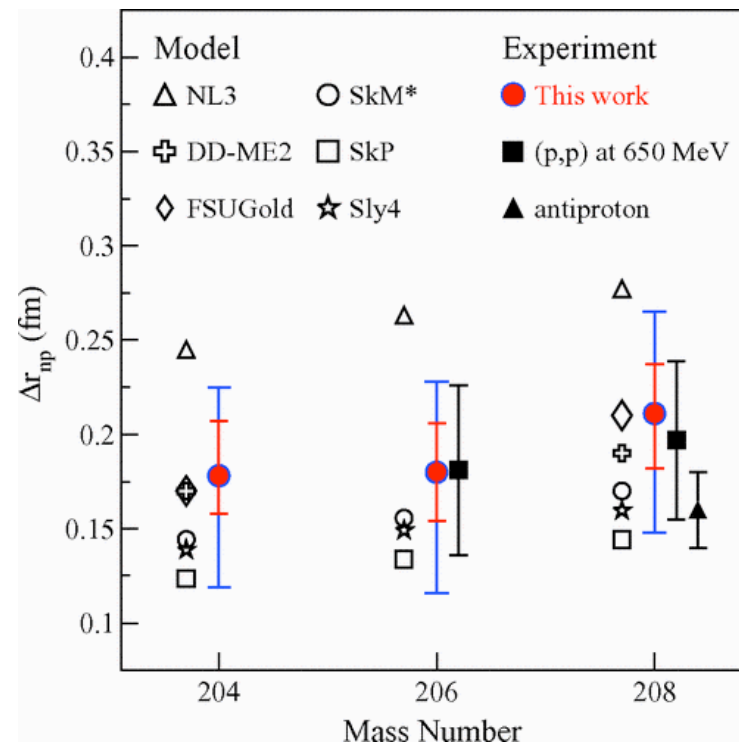
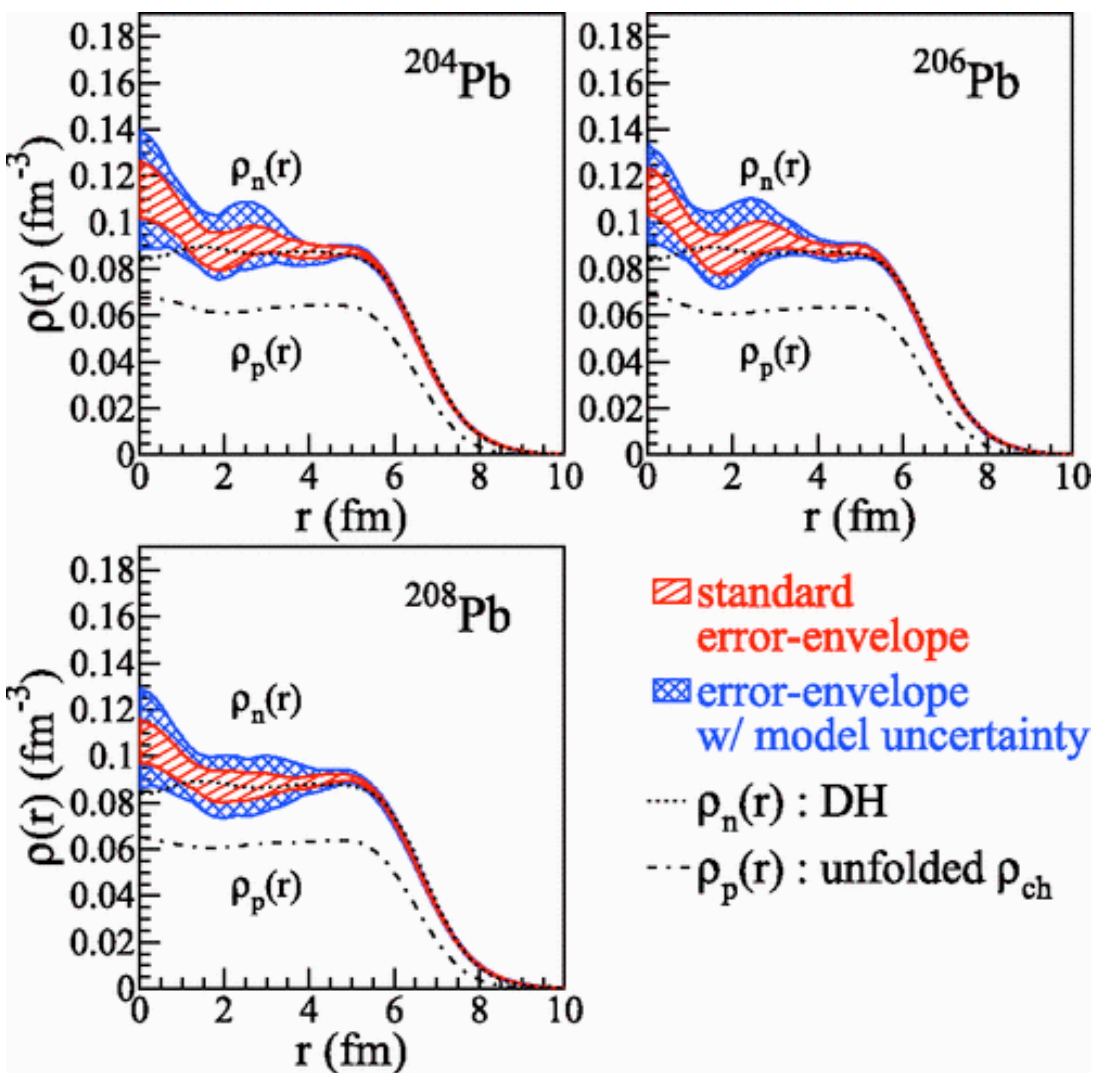
Exp width IAS

~ hundreds of keV

SAMI-ISB: includes ISB  
effects due to Coulomb  
exchange, QED corrections  
and nuclear ISB

# RCNP - proton elastic scattering

Polarized proton elastic scattering @ 295 MeV sensitive to the overall strong size of the nucleus



$$\Delta r_{\text{np}} = 0.211^{+0.054}_{-0.063} \text{ fm}$$

Neutron density distributions of  $^{204,206,208}\text{Pb}$  deduced via proton elastic scattering at  $E_p = 295$  MeV

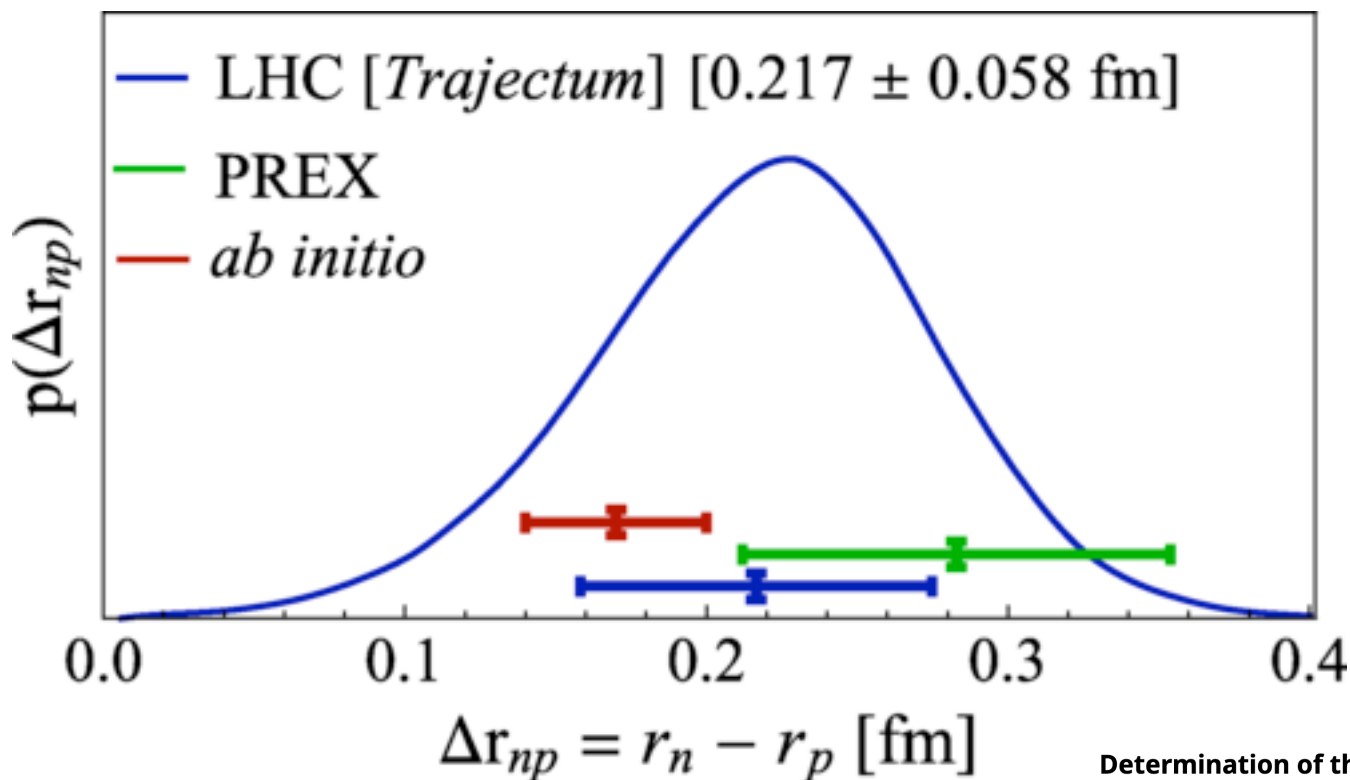
J. Zenhiro<sup>1\*</sup>, H. Sakaguchi<sup>1,2</sup>, T. Murakami<sup>1</sup>, M. Yosoi<sup>1,2</sup>, Y. Yasuda<sup>1,2</sup>, S. Terashima<sup>1,4</sup>, Y. Iwao<sup>1</sup>, H. Takeda<sup>2</sup>, M. Itoh<sup>3,5</sup> et al.

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Phys. Rev. C 82, 044611 - Published 22 October, 2010

# LHC - Relativistic Heavy Ion Collisions

Particle distributions and **collective flow** in  $^{208}\text{Pb} + ^{208}\text{Pb}$  RHIC @ LHC are **sensitive** to the overall strong **size** of the nuclei involved.



Determination of the Neutron Skin of  $^{208}\text{Pb}$  from Ultrarelativistic Nuclear Collisions

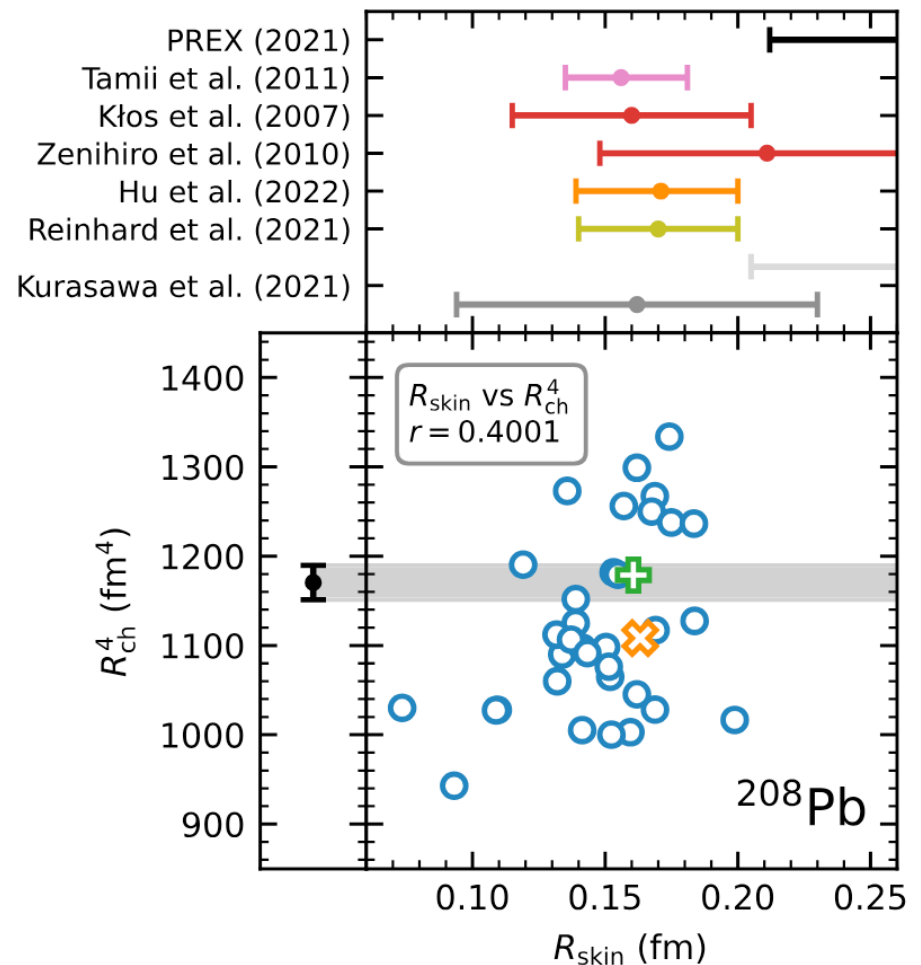
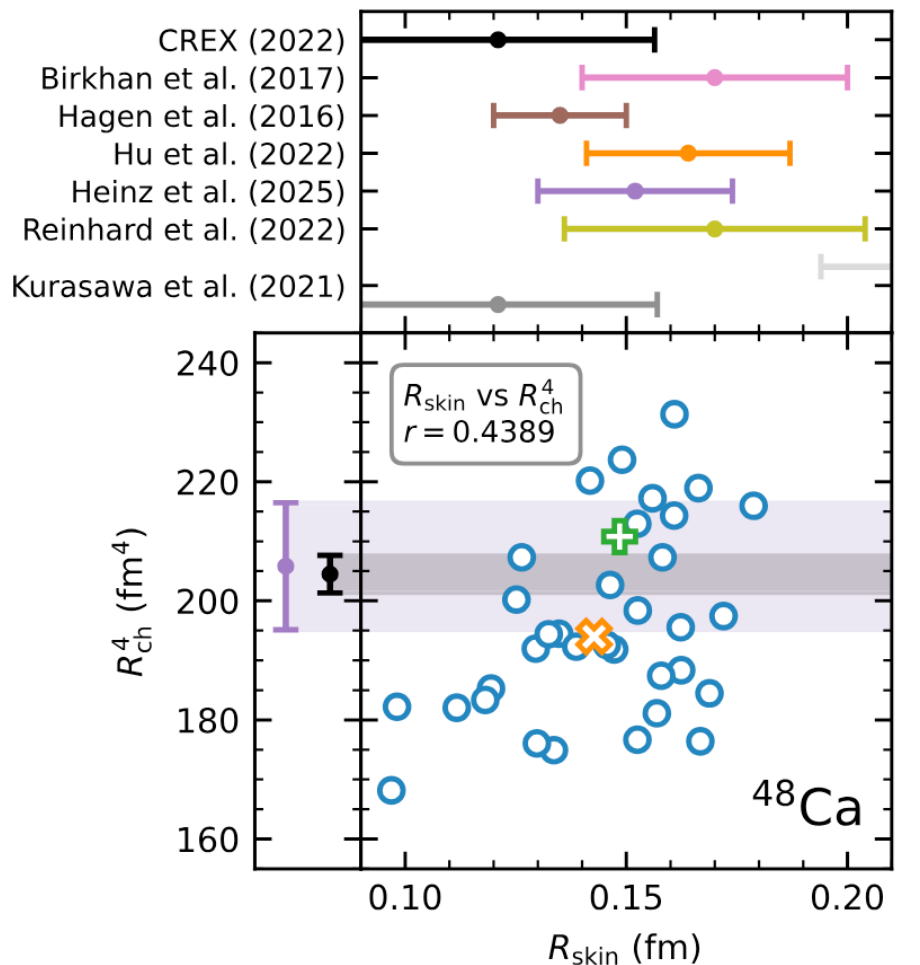
Giuliano Giacalone<sup>1</sup>, Govert Nijss<sup>2</sup>, and Wilke van der Schee<sup>3,4</sup>

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Phys. Rev. Lett. 131, 202302 - Published 15 November, 2023

$$\Delta r_{np} = 0.217 \pm 0.058 \text{ fm}$$

# Summary on $\Delta r_{np}$ in $^{48}\text{Ca}$ and $^{208}\text{Pb}$



$R_{\text{ch}}^4$  inferred from expt.  $R_{\text{ch}}^2$   
 $R_{\text{ch}}^4$  inferred from theor.  $R_{\text{ch}}^2$

$\Delta\text{NNLO}_{\text{GO}}$   
 1.8/2.0 (EM7.5)

Nonimplausible Hamiltonians

Ab initio computations of the fourth-order charge density moments of  $^{48}\text{Ca}$  and  $^{208}\text{Pb}$

# Summary

**Different ways** to learn about the **neutron distribution** in atomic nuclei using **different observables** provide **different answers**

However, within **1.5–2  $\sigma$**  all looks fine, both from phenomenologic **Energy Density Functionals** and from **Ab Initio**

## From Theory:

- An effort to better understand the **parity violating asymmetry** and the **beam normal spin asymmetry**[\*] in  **$^{208}\text{Pb}$**  is needed [PREX & CREX, PRL 128, 142501 (2022)].
- An effort to better understand the **systematics** on the **dipole polarizability** would be of interest [see e.g. Sn chain in Bassauer et al. PLB 810, 135804 (2020)].

## From Experiment (low-energy):

- An effort to improve the accuracy in the **parity violating asymmetry in  $^{208}\text{Pb}$**  (and measure **other Q values**) **is needed**. Measuring other nuclei would also be desirable.
- **Systematic** measurements of the **dipole polarizability** along **neutron rich** isotopic chains (e.g.  $N > 74$  Sn isotopes) could help testing models and improve our understanding of this observable.

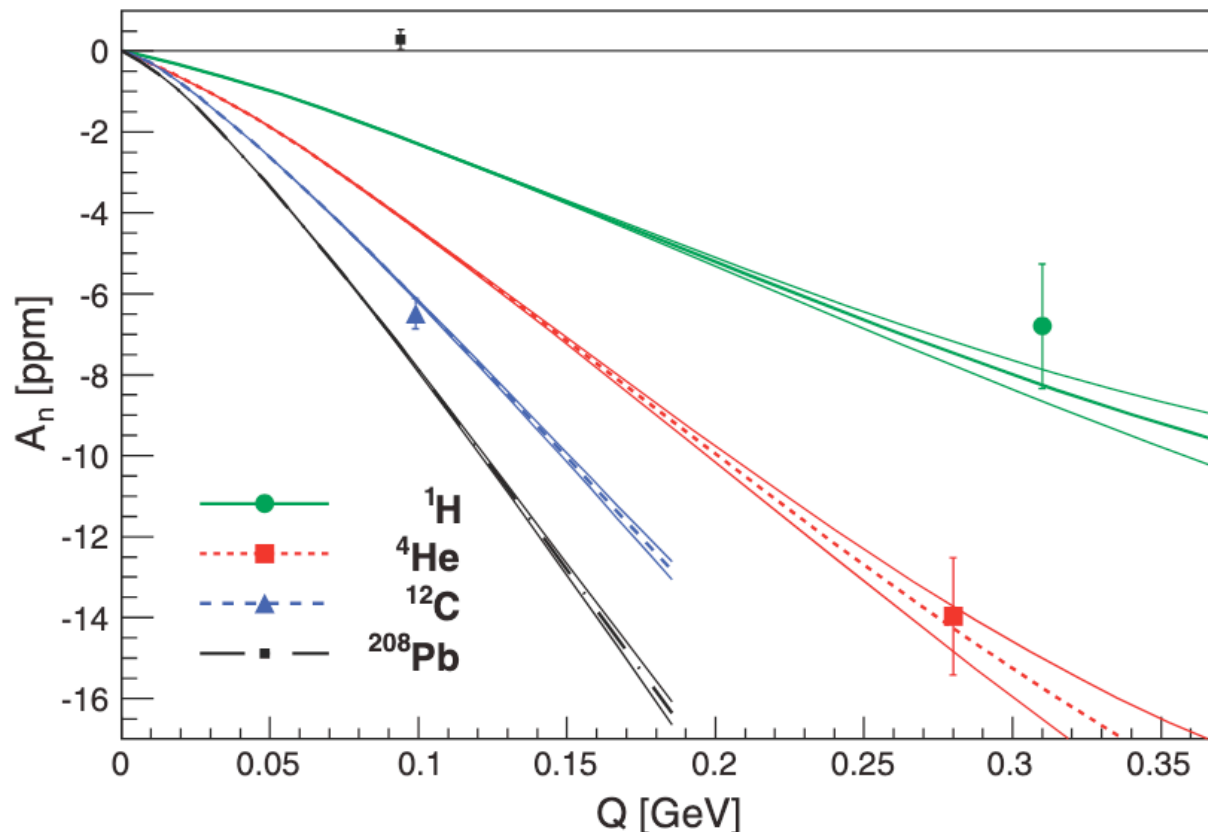


# Collaborators

- Gianluca **Colò** (University of Milan)
- Pietro **Klausner** (University of Milan)
- Tomoya **Naito** (University of Tokyo)
- Xavier **Vinyes** & Mario **Centelles** (University of Barcelona)
- Jorge **Piekarewicz** (Florida State University)
- Nils **Paar** & Dario **Vretenar** (University of Zagreb)
- Bijay K. **Agrawal** (Saha Institute of Nuclear Physics)
- P.-G. **Reinhard** (University of Erlangen-Nürnberg)
- Witold **Nazarewicz** (FRIB and Michigan State University)
- Doris H. **Jakubassa-Amunsden** (Ludwig-Maximilians-Universität München)

# Beam normal spin asymmetry ( $A_n$ ) (aka Alaying power or Sherman function)

Is the asymmetry in the scattering cross section when the incident electron beam is polarized perpendicular (longitudinal  $\rightarrow A_{pv}$ ) to the scattering plane



$\Rightarrow$  Important to understand  $A_{pv}$  since it represents a potentially large systematic correction to measured asymmetries.

In Born approximation (one-photon exchange),  $A_n = 0$  because time-reversal symmetry forbids such an asymmetry.

Arises from interference between one-photon exchange and the absorptive part of two-photon exchange amplitudes.

## New Measurements of the Transverse Beam Asymmetry for Elastic Electron Scattering from Selected Nuclei

[S. Abrahamyan](#)<sup>45</sup>, [A. Acha](#)<sup>10</sup>, [A. Afanasev](#)<sup>11</sup>, [Z. Ahmed](#)<sup>33</sup>, [H. Albataineh](#)<sup>6</sup>, [K. Aniol](#)<sup>3</sup>, [D. S. Armstrong](#)<sup>7</sup>, [W. Armstrong](#)<sup>35</sup>, [J. Arrington](#)<sup>1</sup> et al. (HAPPEX and PREX Collaborations)

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Phys. Rev. Lett. **109**, 192501 – Published 5 November, 2012  
DOI: <https://doi.org/10.1103/PhysRevLett.109.192501>

## New Measurements of the Beam-Normal Single Spin Asymmetry in Elastic Electron Scattering over a Range of Spin-0 Nuclei

[D. Adhikari](#)<sup>1</sup>, [H. Albataineh](#)<sup>2</sup>, [D. Androic](#)<sup>3</sup>, [K. Aniol](#)<sup>4</sup>, [D. S. Armstrong](#)<sup>5</sup>, [T. Averett](#)<sup>5</sup>, [C. Ayerbe Gayoso](#)<sup>5</sup>, [S. Barcus](#)<sup>5</sup>, [V. Bellini](#)<sup>7</sup> et al. (PREX and CREX Collaborations)

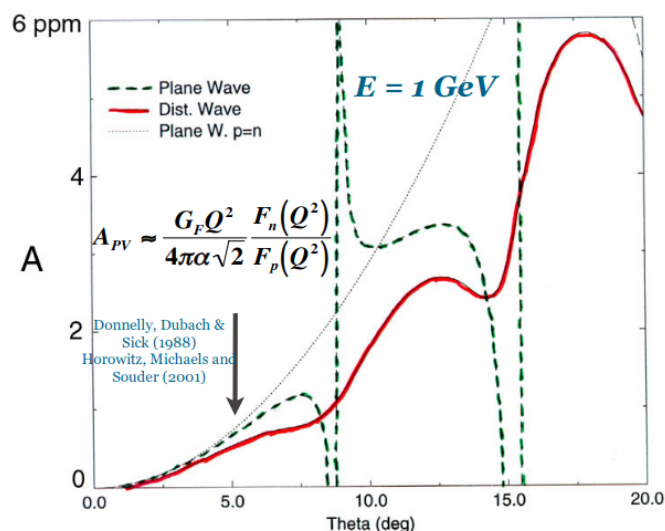
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Phys. Rev. Lett. **128**, 142501 – Published 8 April, 2022  
DOI: <https://doi.org/10.1103/PhysRevLett.128.142501>

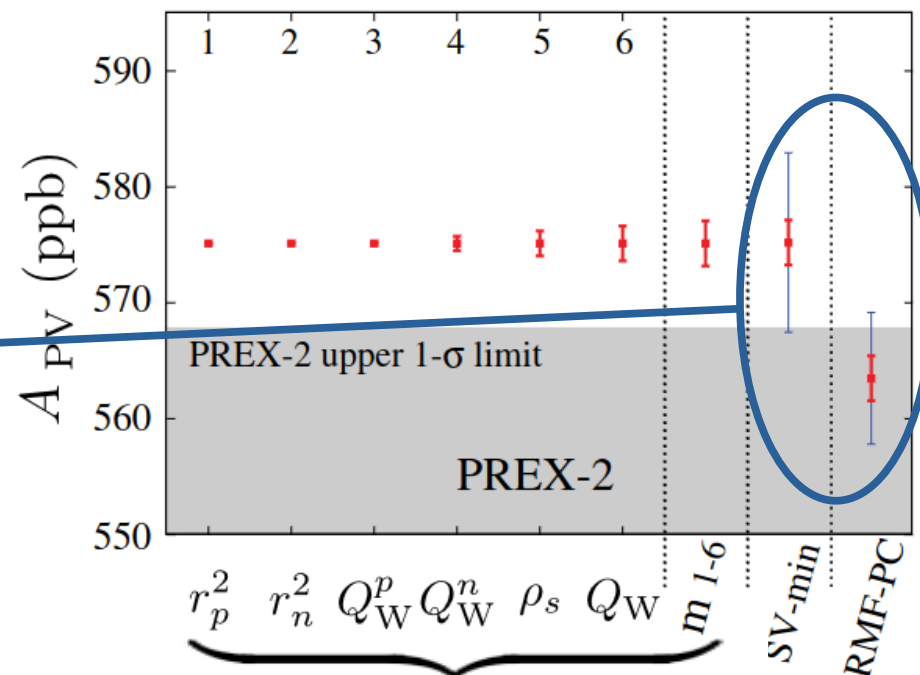
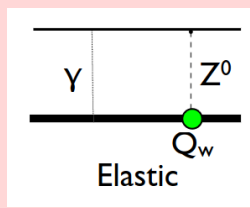
# Parity Violating Asimmetry: theory

The main uncertainties come from nuclear model errors on the description of the neutron distribution (blue error bars)

→ Coulomb distortions are important (order  $\alpha Z$ )



Are  $\gamma Z_0$  box corrections important? not included

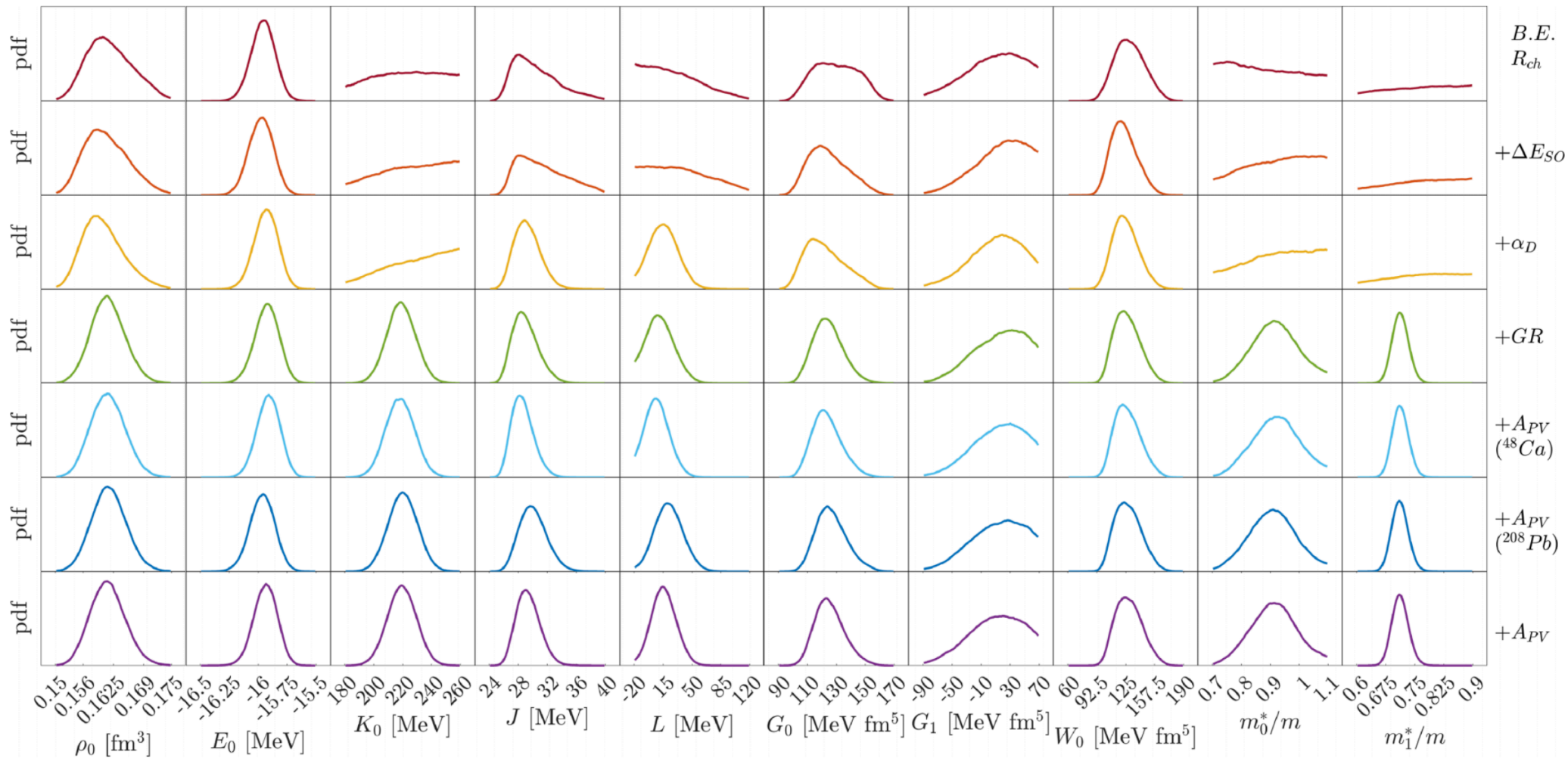


Hadronic uncertainties less relevant

| Parameter                                  | Value                |
|--|----------------------|
| $\langle r_p^2 \rangle$ (fm <sup>2</sup> ) | $0.726 \pm 0.019$    |
| $\langle r_n^2 \rangle$ (fm <sup>2</sup> ) | $-0.1161 \pm 0.0022$ |
| $\mu_p$                                    | 2.792 847            |
| $\mu_n$                                    | -1.9130              |
| $Q_p^{(W)}$                                | $0.0713 \pm 0.0001$  |
| $Q_n^{(W)}$                                | $-0.9888 \pm 0.0011$ |
| $\rho_s$                                   | $-0.24 \pm 0.70$     |
| $\kappa_s$                                 | $-0.017 \pm 0.004$   |
| $Q_{126,82}^{(W)}$                         | $-117.9 \pm 0.3$     |

# Bayesian inference Skyrme EDF: can we accommodate $\alpha_D$ and $A_{PV}$ without compromising other observables?

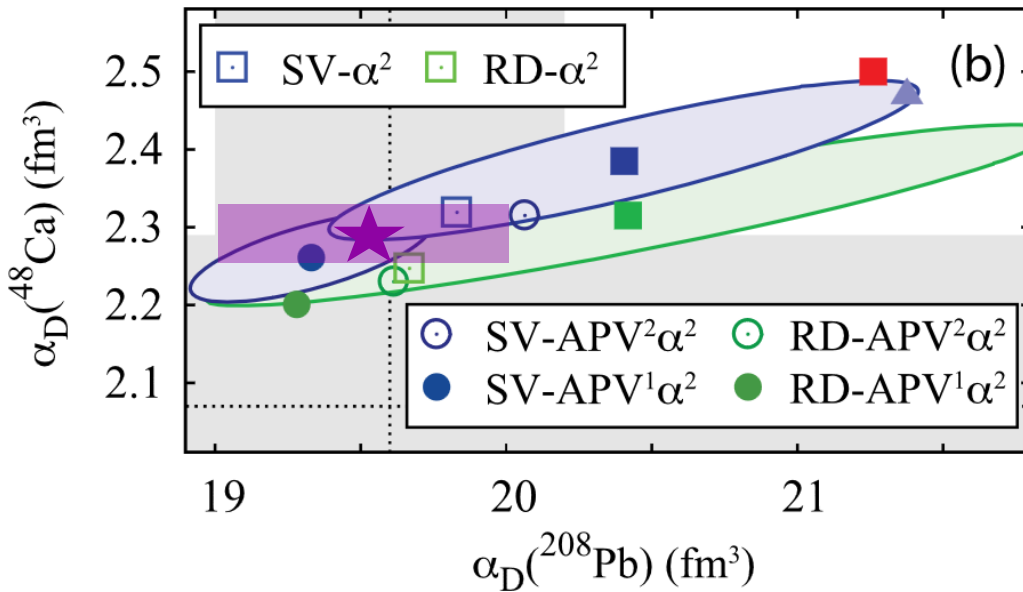
By P. Klausner



# Bayesian inference Skyrme EDF: can we accommodate $\alpha_D$ and $A_{PV}$ without compromising other observables?

Pietro Klausner, Gianluca Colò, Xavier Roca-Maza, and Enrico Vigezzi  
Phys. Rev. C 111 014311 (2025)

By P. Klausner

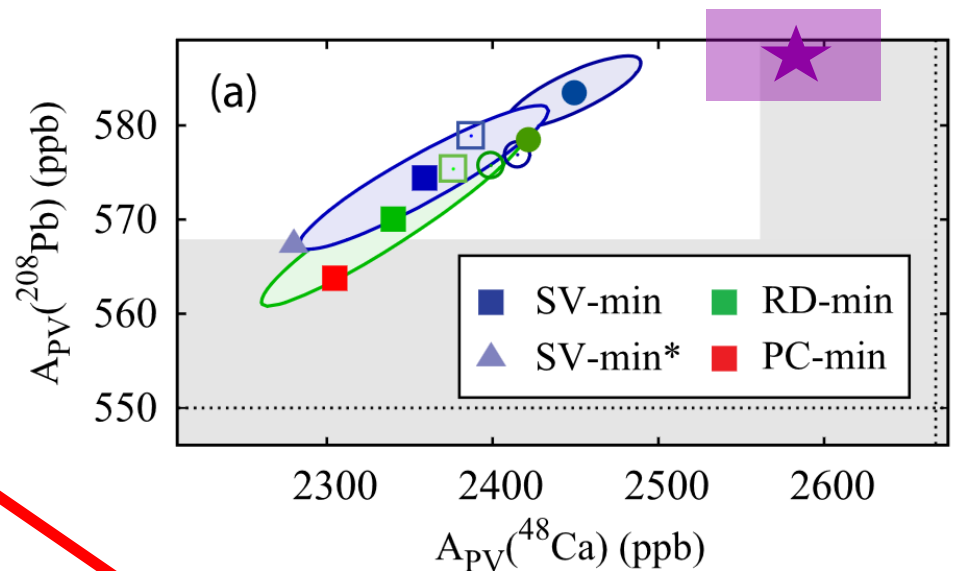


|                   | $\alpha_D$ (fm <sup>3</sup> ) | $m(1)$ (MeV fm <sup>2</sup> ) | $A_{PV}$ (ppb) |
|-------------------|-------------------------------|-------------------------------|----------------|
| <sup>208</sup> Pb | $19.5 \pm 0.5$                | $958 \pm 22$                  | $589 \pm 5$    |
| <sup>48</sup> Ca  | $2.30 \pm 0.08$               | —                             | $2591 \pm 54$  |

|                   | B.E. (MeV)     | $R_{ch}$ (fm)   | $\Delta E_{SO}$ (MeV) |
|-------------------|----------------|-----------------|-----------------------|
| <sup>208</sup> Pb | $1636 \pm 1.8$ | $5.49 \pm 0.03$ | $2.34 \pm 0.16$       |
| <sup>48</sup> Ca  | $417 \pm 1.2$  | $3.51 \pm 0.02$ | $1.92 \pm 0.20$       |
| <sup>40</sup> Ca  | $342 \pm 1.6$  | $3.50 \pm 0.02$ | —                     |
| <sup>56</sup> Ni  | $482 \pm 1.4$  | —               | —                     |
| <sup>68</sup> Ni  | $590 \pm 1.0$  | —               | —                     |
| <sup>100</sup> Sn | $826 \pm 1.6$  | —               | —                     |
| <sup>132</sup> Sn | $1103 \pm 1.7$ | $4.71 \pm 0.03$ | —                     |
| <sup>90</sup> Zr  | $784 \pm 1.3$  | $4.27 \pm 0.02$ | —                     |

Isoscalar resonances

|                   | $E_{GMR}^{IS}$ (MeV) | $E_{GQR}^{IS}$ (MeV) |
|-------------------|----------------------|----------------------|
| <sup>208</sup> Pb | $13.5 \pm 0.3$       | $10.8 \pm 0.4$       |
| <sup>90</sup> Zr  | $17.8 \pm 0.4$       | —                    |



Keeping ground and excited state properties within typical Skyrme-EDF accuracy